

EANAM10, September 15-19, 2025 <https://coma.kasi.re.kr/eanam10/program/>

# Multi-component-abundance-isotope SEPs acceleration by coupled hydrodynamic, magnetodynamic, and kinetic processes in fractal current sheets in solar flare/CMES

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Sciences, State Key Laboratory of Space Weather, National Space Science Center, CAS  
Yunnan Key Laboratory of Solar Physics and Space Science

## Cooperate with

**Yongbing Li**, Yang Zhao, Qian Wang: Earth and Planetary College, University of Chinese Academy of Sciences  
Yan Li, Zhikuo Ma, Zehua Fu, Zhihe Wang: Yunnan Observatories; Yunnan Key Laboratory of Solar Physics and Space Science.

**Hui Yan**, Ying Zhong: National Supercomputer Center in Guangzhou, Sun Yat-sen University ([Tianhe-2 CPU](#))

**Wu Wang**, Yufeng Guo, Qing Ji: National Supercomputer Center in CAS, CNIC ([Dongfang series, CPU-DCU](#))

**Chunye Gong**, Gang Zheng, Chuan Wang: National Supercomputer Center in Tianjin & National University of Defense Technology ([CPU-ARM](#))

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(CU Boulder) George Ignatius Fann (OKA), Robert von Fay-Siebenburgen(UOS), Hua-Liang Wei(UOS), Matthew Gregg Knepley (Mark F Adams(LBNL), Hui Li(LANL), Shengtai Li(LANL), **David A Yuen(CU)** et.al

15:00-15:20 SEP 17, 2025



# OUTLINE

## 1. Background and Motivation

What & Why, Research progress, Science points, How- Key points

## 2. RHPIC-LBM: X86CPU & CPU-GPU & CPU-DCU & ARM CPU

Method, Algorithm, code, validation

## 3. Application (Events Study)

3D turbulence acceleration[source and origin, Electron, Proton,  $^3\text{He}$ ,  $^4\text{He}$ ;  $^3\text{He}$  rich]

## 4. Summary and Future Works

Traditional turbulence acceleration(1st & 2nd Fermi) and wave-particle interaction acceleration  
Other heavy particles & Different ionization states

# 1. Background and Motivation

**What & Why:** GeV-level-SEPs space weather disasters vs. source-origin(Heavy particle,  $^3\text{He}$  rich)

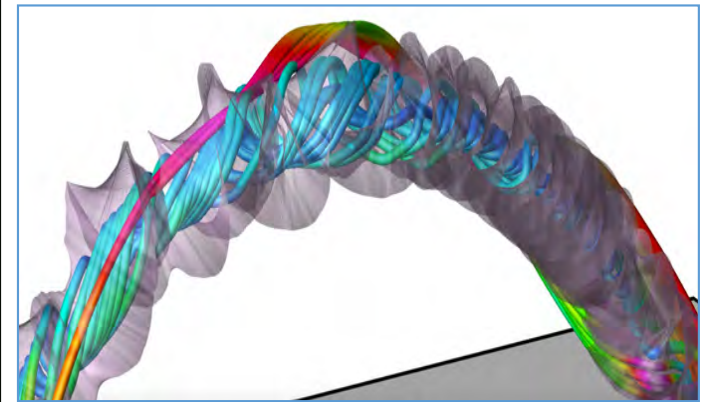
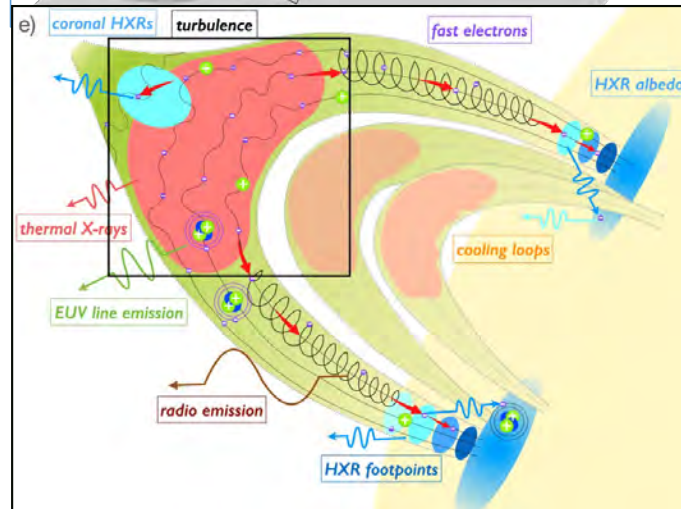
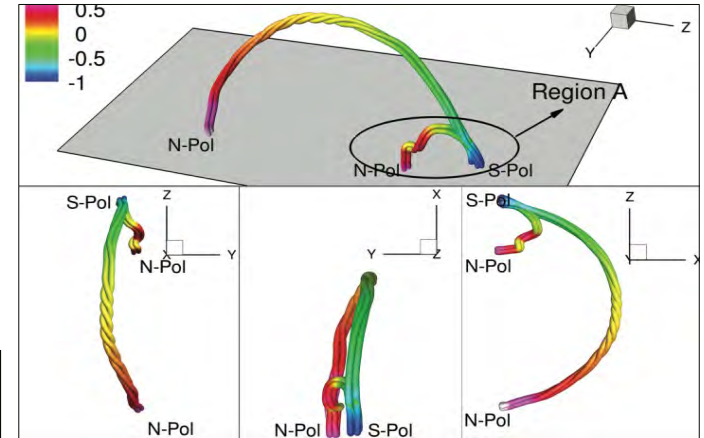
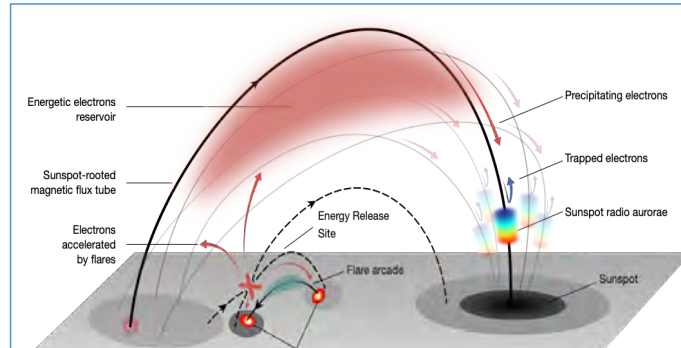
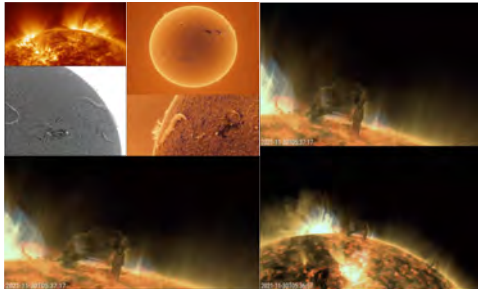
**Research progress:** Supercomputer(SUGON,TIANHE-3) & observational data(FY-3E heavy SEPs)

**How- Science points:** Multi-component-abundance-isotope, Turbulence induced viscosity & resistivity dissipation and diffusion

**Key points:** 3D turbulence MR, 3D topologies evolution, Fractal CS mode

# Motivation: SEPs induced extremely space weather disasters

GeV-level-SEPs-induced extreme space weather disasters. **identifying producing source and origin**



# Research progress : Supercomputer & Observational data

Our research team: **Early warning system of SEPS-induced space weather disasters** with the observational data-driven multi-component-abundance-isotope model through plasma statistical physics theoretical framework on the domestic supercomputer

## FY-3E SEPs Observational data

中国气象数据网 | 首页 | 数据服务 | 接口服务 | 数据汇交 | 可视化 | 气象专题 | 服务案例 | 用户支持 | 关于我们 | 北京 14.0°C

气象搜索 高能粒子

返回搜索 再分析

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朱伯靖 **数据科学家**  
等级: 7级  
注册日期: 2014-05-27 10:00:00  
积分: 100

SEM高能粒子产品

SEM高能粒子产品

属于FY3C卫星SEM仪器, SEM高能粒子和电位产品提供了空间环境监测器在卫星轨道高度上探测到的高能质子、电子和重离子 (He、Li、C、Mg、Fe) 的辐射通量, 每天14轨数据。产品内容包括时间信息、高度、地理经纬度、地磁经纬度、磁壳L值、6道质子通量、离子谱计、5道电子通量、重离子信息 (H, Li, C, Mg, Ar, Fe)

数据名称 SEM高能粒子产品

关键字 FY3C FY3c FY-3C FY-3e SEM Sem 空间环境监测器 空间环境监测器 SPACE ENVIRONMENT MONITOR Space Environment Monitor EPP Epo 高能粒子 高能粒子 HIGH ENERGY PARTICLE PRODUCT High Energy Particle Product  
FY3C\_SEMWX\_ORBIT\_L2\_EPP\_MLT\_NUL\_VVVVMMDD\_HHMM\_00000\_MS.DAT  
FY3c\_semw\_orbit\_l2\_app\_mlt\_nul\_vvvvmmdd\_hhmm\_00000\_ms.dat SEM高能粒子产品 Sem高能粒子产品 SEM HIGH ENERGY PARTICLE PRODUCT Sem High Energy Particle Product

空间范围 全球

制作时间 实时

数据起始时间 2014-05-27 00:00:00

SEM高能粒子产品 (下列检索条件均需填写, 其中时间条件为世界时)

日期选择 2024-05-13 00:00:00 至 2024-05-13 24:00:00

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搜索



# Science points: Multi-components, multi-abundance, multi-isotope

Journal Pre-proofs

Review of Solar Energetic Particle Models

Kathryn Whitman, Ricky Egeland, Ian G. Richardson, Clayton Allison, Philip Quinn, Janet Barzilla, Inna Kisilashvili, Viacheslav Sadykov, Hazel M. Bain, Mark Dieckmann, M. Lella Mays, Tilaye Tadese, Kerry T. Lee, Edward Semones, Janet G. Luhmann, Marlon Néñez, Stephen M. White, Stephen W. Kabler, Alan G. Ling, Don F. Smart, Margaret A. Shea, Valery Tenishev, Soukaina F. Boubrabani, Berkay Aydin, Petrus Martens, Rafal Angryk, Michael S. Marsh, Silvia Dalla, Norma Crosby, Nathan A. Schwadron, Kamen Kozarev, Matthew Goebel, Matthew A. Young, Monica Laurenza, Edward W. Cliver, Tommaso Alberti, Mirko Stumpo, Simone Benelli, Athanasios Pappasou, Anastasios Anastasiadis, Ingar Sandberg, Manolis K. Georgoulis, Anli Ji, Dustin Kempton, Chetraj Pandey, Gang Li, Junxiang Hu, Gary P. Zank, Eleni Lavasa, Giorgos...

Yash Kadadi, Ian Fernandes, Maher A. Di Subbanyo, Chatterjee, Kimberly D. Morel, Roussev, Aleksandr Takakhivili, Freda Zhengping Huang, Lulu Zhao, Nicolas W. Athanasios Kouloumvakos, Mikka Passilita, Rami Vainio, Anatoly Belov, Eugenia A. Eroshenko, Maria A. Abunin, Artem A. Abunin, Christopher C. Balch, Olga Malandraki, Michalis Karavolos, Bernd Heber, Johannes Labrenz, Patrick Kühl, Alexander G. Kosovichev, Vincent Oria, Gela M. Nita, Igor Ilarionov, Patrick M. O'Keefe, Yucheng Jiang, Sheldon H. Freese, Ainyi Ali, Evangelos Paschos, Sigrav Annalagaria-Giamini, Piers Jeggens, Meng Jin, Christina O. Lee, Erika Palmerio, Alessandro Bruno, Spiridon Kasapi, Xiantong Wang, Yang Chen, Bhai Sanubaja, David Lario, Carla Jacobs, Du Tott Straus, Ruhana Steyn, Jabos den Berg van, Bill Swallow, Charlotte Waterfall, Mohamed Nodal, Rosina Miteva, Monchal Dechev, Pietro Zucca, Alec Engel, Brianna Maze, Harsh Farmer, Thaha Kerber, Ben Barnett, Jeremy Loomis, Nathan Grey, Barbara J. Thompson, Jon A. Linker, Ronald M. Cegles, Cooper Downs, Tibor Tóth, Roberto Lionello, Viacheslav Titov, Ming Zhang, Pooya Hossainzadeh

PII: S0273-1177(22)00724-4  
DOI: <https://doi.org/10.1016/j.asr.2022.08.006>  
Reference: JASR 16144



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## Traditional turbulence acceleration(1st & 2nd Fermi) and wave-particle interaction acceleration

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### Donald Reames

(<https://ipst.umd.edu/people/donald-reames>)

Solar Energetic Particles: Spatial Extent and Implications of the H and He Abundances

Donald W. Reames

Received 10 May 2022 | Accepted 7 July 2022 | Published online 21 August 2022

65th Years of Element Abundance Measurements in Solar Energetic Particles

Donald W. Reames

**A Perspective on Solar Energetic Particles**

Abstract: This review discusses the current status of solar energetic particle (SEP) science, with a focus on the H and He abundances. It covers the historical context, the current state of knowledge, and the challenges ahead. The review is organized into several sections: Introduction, Solar Energetic Particles: Spatial Extent and Implications of the H and He Abundances, and A Perspective on Solar Energetic Particles.

**INTRODUCTION**

The study of solar energetic particles (SEPs) is a multidisciplinary field that spans across several areas of solar and space physics. It is a field that has seen rapid growth in the past few decades, and it is expected to continue to do so in the future. This review provides a perspective on the current state of SEP science, with a focus on the H and He abundances.

**NUCLEAR EMULSION**

The first measurements of SEP abundances were made using nuclear emulsion detectors. These detectors were used to measure the H and He abundances of SEPs, and they provided the first evidence for the existence of SEPs. The use of nuclear emulsion detectors was limited by their low efficiency and their inability to measure the energy of the particles. However, they were the only method available at the time for measuring the H and He abundances of SEPs.

**Keywords:** Solar energetic particles, Shock waves, Coronal mass ejections, Solar jets, Solar flares, Solar system abundances, Solar wind

**1 Introduction**

The evidence for solar energetic particles begins with the largest events, where GeV protons and nuclear cascades of particles through the Earth's atmosphere that were observed by air showbers at ground level by Forbush (1944). The solar energetic particles (SEPs) in these ground-level events (GLEs) were shown to be continuous background produced in the same manner by the galactic cosmic rays (GCRs), and the protons of the SEPs provided broad evidence in the GCR, namely, a third of the flux, that was understood to be caused by scattering of the GCRs by magnetic fields of coronal mass ejections (CMEs) originating from the Sun. This is now well established by Kallen et al. (1995) showed a 90% correlation between SEP and these mass ejections, and CMEs that drive shock waves

Table 1. Chemical composition of the Sun

Chemical composition	Percent of the overall mass	Type
<b>Hydrogen</b>	<b>73%</b>	<b>Dominated</b>
Helium	25%	Dominated
Oxygen	0.8%	Metals
Carbon	0.36%	Metals
Iron	0.16%	Metals
Neon	0.12%	
Nitrogen	0.09%	
Silicon	0.07%	Metals
Magnesium	0.05%	Metals
Sulphur	0.04%	Metals
Others combined	0.04%	Metals

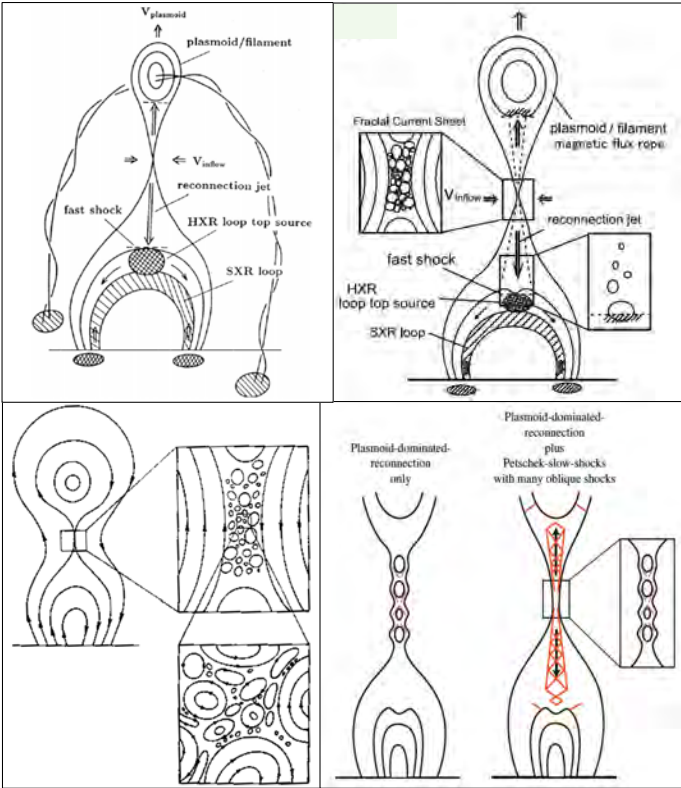
## Other heavy particles & Different ionization states

Table 2. Layers of the Sun

Layers	Thickness (km)	Temperature (K)	Type
Core	150,000	10,000,000	Inertial layer
Convection Zone	180,000	500,000	Inertial layer
		0	Outer layer, un-magnetized, weakly ionized
		8000	Outer layer, magnetized, partly ionized <sup>2</sup>
Transition region	100	8000 – 500,000	Outer layer, magnetized, high ionized
Corona	5,000,000	1,000,000	Outer layer, magnetized, fully ionized

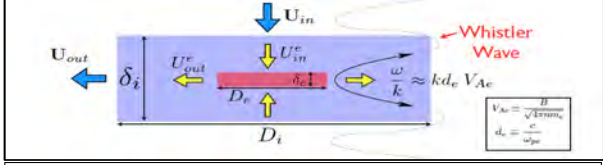
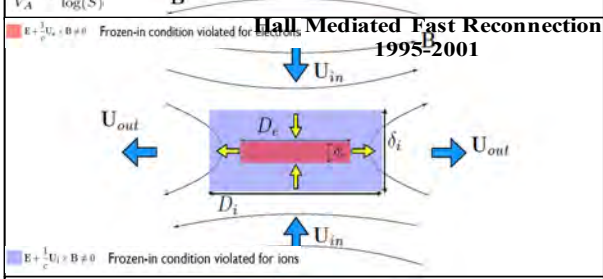
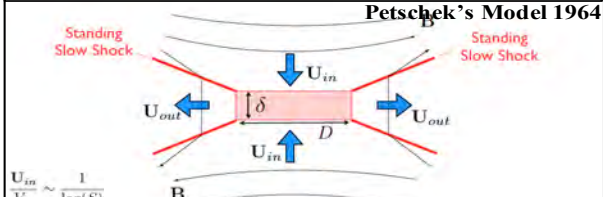
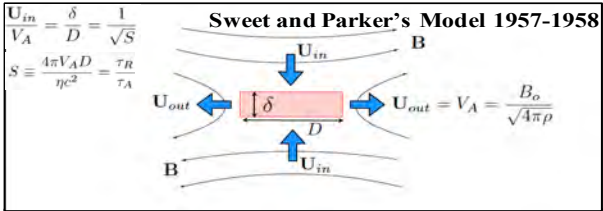
# Science points: Fractal CS model

B&U decoupled condition and Fine structure investigation  
 Turbulence induced viscosity & resistivity dissipation and diffusion

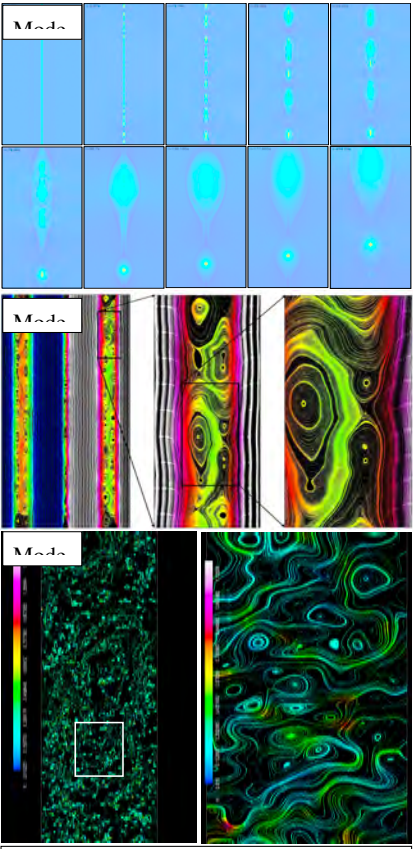


Shibata et al., 1995; Tsuneta et al., 1992; Masuda et al., 1994; Tajima & Shibata 1997; Mann et al., 2006; Zenitani and Miyoshi 2011; Zenitani 2015

Theoretical Model from Observation



Theoretical Model from Plasma



Numerical Model from Simulation

Model A: **single current sheet?**

Zenitani and Miyoshi (2011) and Zenitani (2015)

Model B: **several current sheet?**

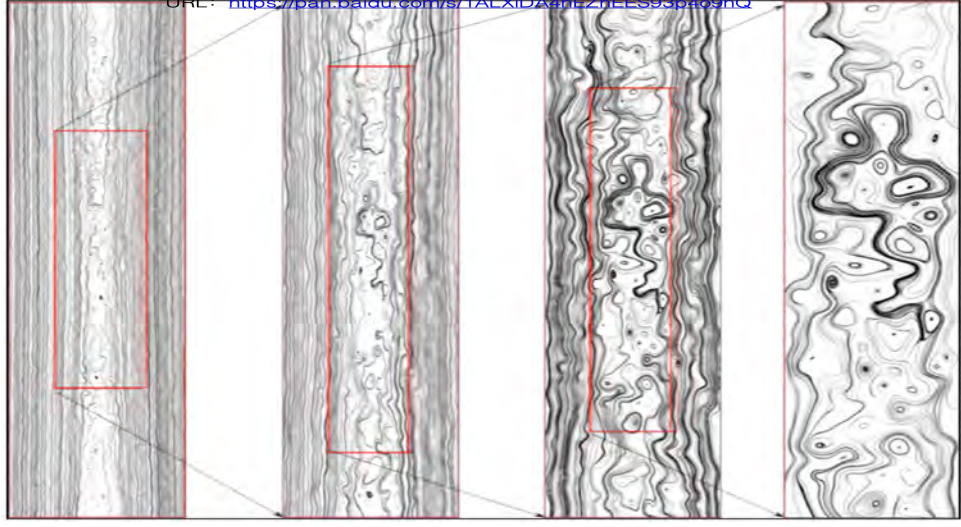
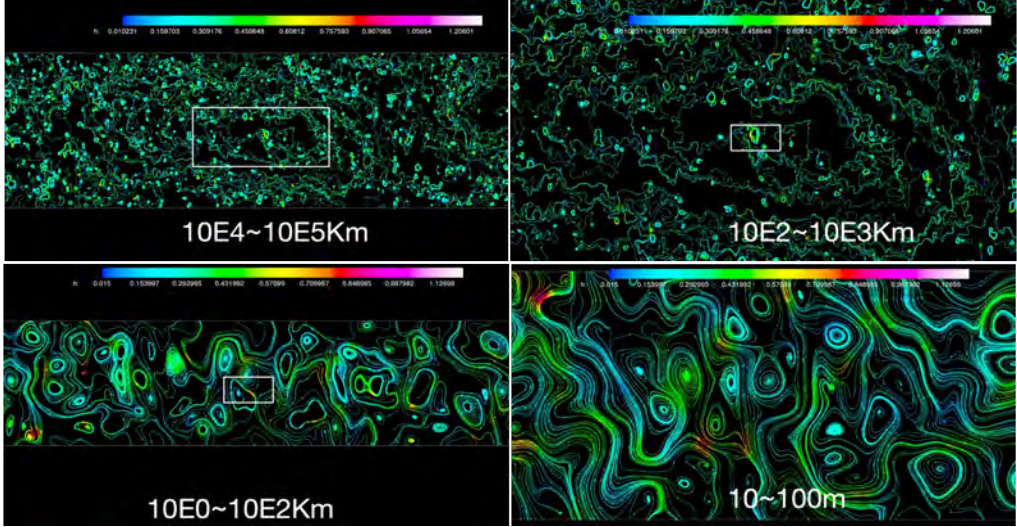
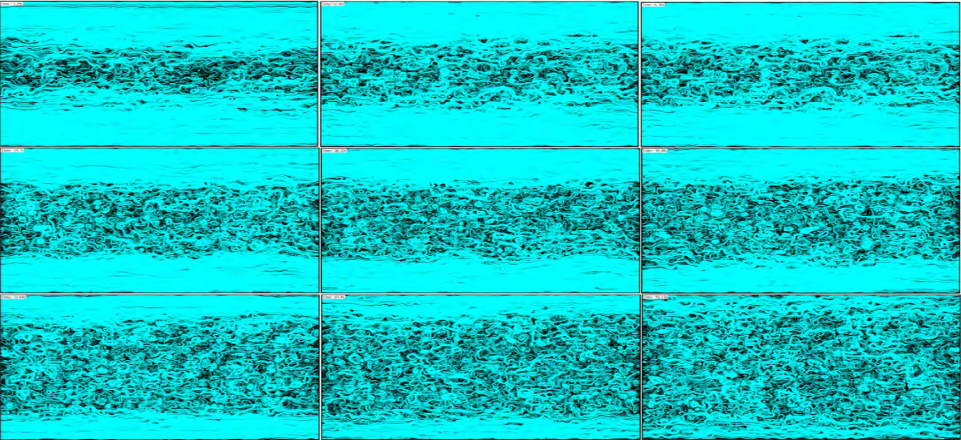
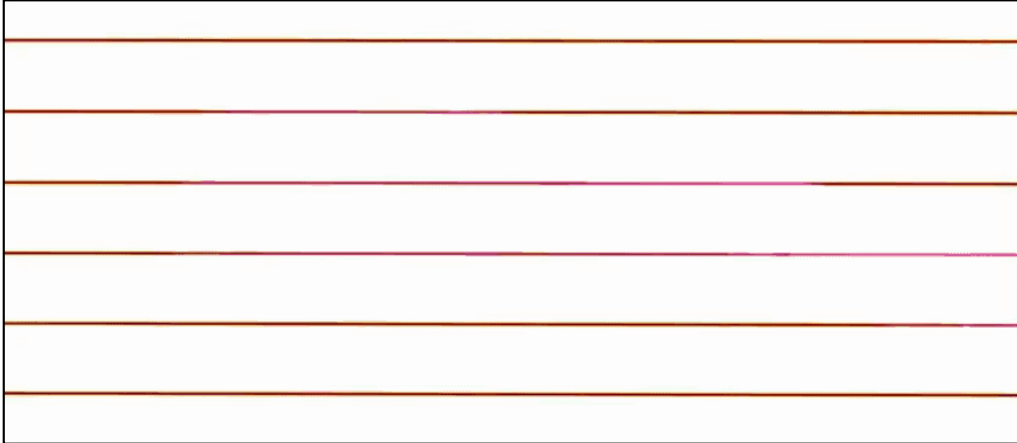
D. Burgess, P. W. Gingell, and L. Matteini (2016)

Model C: **full of kinetic scale current sheet?**

Lazarian & Vishniac (1999), Servidio et al. (2009, 2010, 2011), G. P. Zank et al (2014)

# Key points: 3D turbulence MR

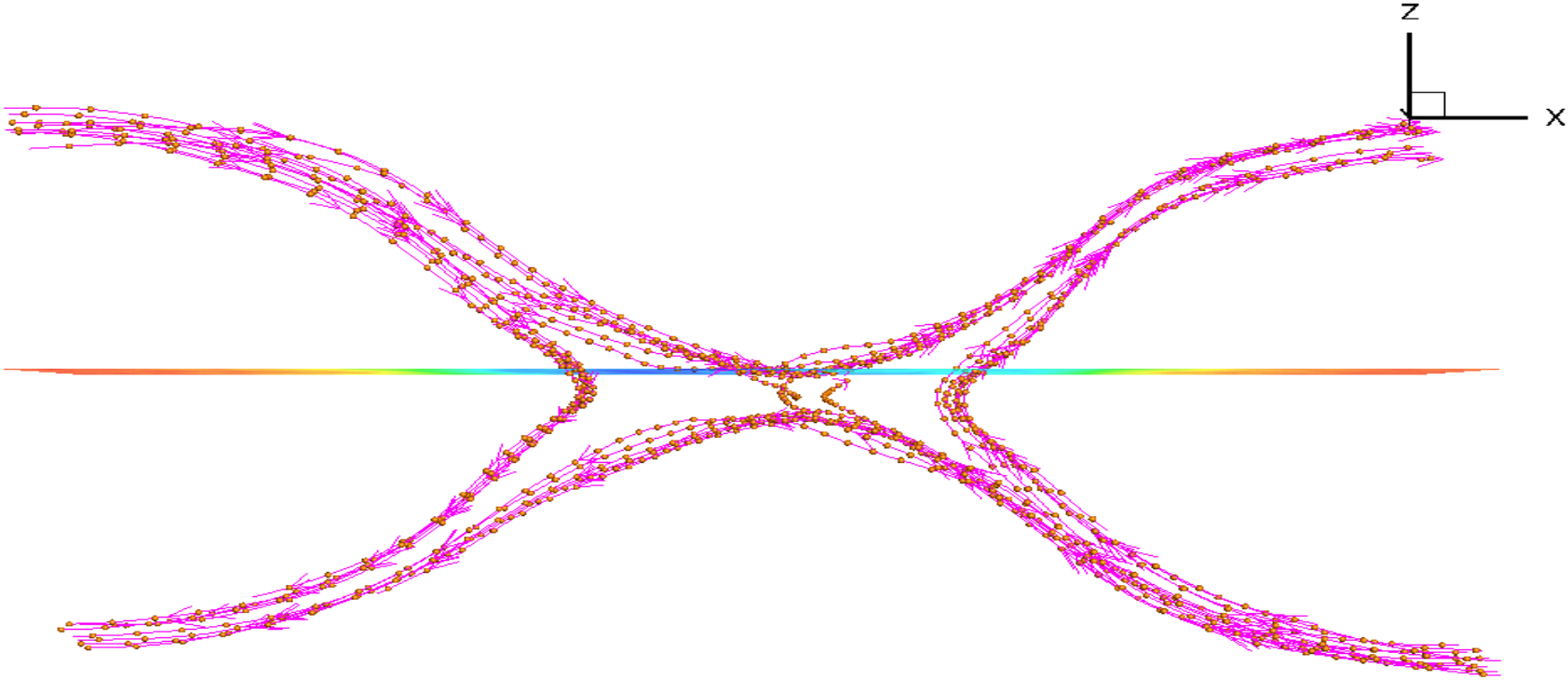
Turbulence plays a key role in the energy diffusion-dispersion-dissipation, flare heating and SEP energization and acceleration.



BOJING ZHU 2016 (Zhu et al 2019,2020a,2020b). Video URL: <https://pan.baidu.com/s/1ALXiDA4hEZhEES93p4o9nQ>

# Key points: 3D turbulence MR

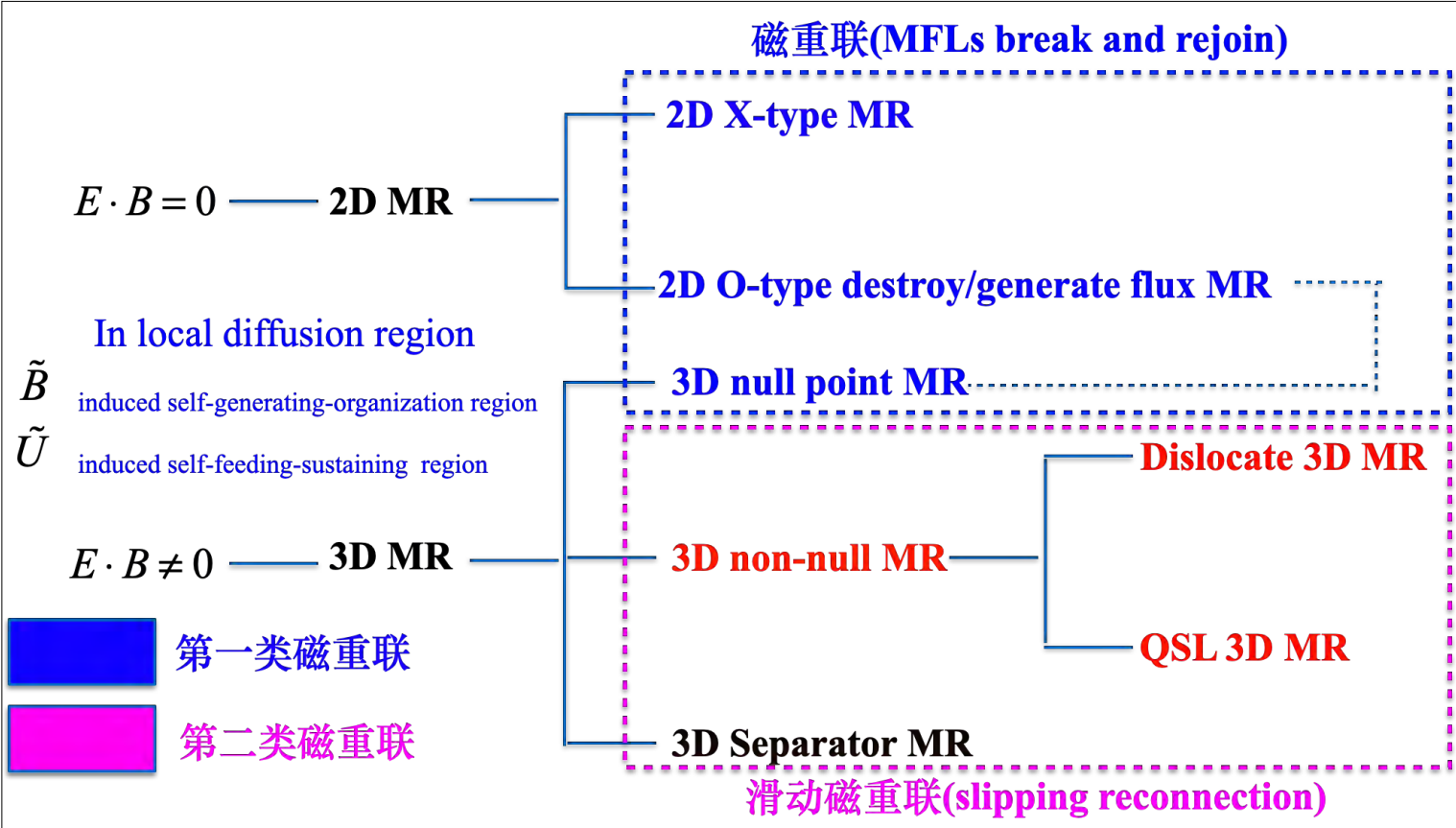
Fundamentally three-dimensional: What is hidden in the interim of the vorticity (plasma bubble; magnetic islands/magnetic helical structure)



Birn and Priest 2007; Pontin 2011; Gonzales and Parker 2016; Zhu et al 2019,2020a,2020b:

# Key points: 3D turbulence MR

Fundamentally three-dimensional: What is hidden in the interim of the vorticity (plasma bubble; magnetic islands/magnetic helical structure)



**Magnetic islands:** defined the **magnetic island** as a kind of magnetic field structure consisting of a bipolar magnetic field normal to the CS, with an enhancement of the core field in the axial direction with a compressed and high-temperature plasma inside of it.

**Flux rope:** Initial small magnetic helical perturbation will change the geometry and topology of the magnetic field, making MFLs orthogonal to the vector of the perturbation  $\tilde{B}$ , forming a magnetic helical structure upon the flux surface (consisting of many spiral MFLs).

## **2. RHPIC-LBM**

Method, Algorithm, Code, Validation

# RHPIC-LBM: Method & Algorithm & code & Validation

**3D Hybrid Hypersingular Integral Equation(HHIE)**  
**3D Hybrid Hypersingular Integral Equation-Lattice Boltzmann Method(HHIE-LBM)**  
**3D Lattice Green Function-Lattice Boltzmann Method(LGM-LBM) Parallel CPU & GPU**

Hybrid method combined the Particle-in-Cell (PIC) and Lattice Boltzmann Method (HPIC-LBM)  
 (2D/2.5D Kinetic-Dynamic-Hydro)

Relativistic HPIC-LBM  
 (2D/2.5D Kinetic-Dynamic-Hydro)

Validation

Developed and added new Program modules: Turbulence induced anomalous resistivity and viscosity  
 (3D Kinetic-Dynamic-Hydro)

**Parallel CPU&GPU Version**  
 (3D Kinetic-Dynamic-Hydro)

2002.09 ~ 2008.01  
 2008.01 ~ 2013.01  
 2013.01 ~ 2015.07  
 2015.07 ~ 2016.10  
 2016.10 ~ 2017.10  
 2017.10 ~ 2018.03  
 2018.03 ~ 2021.07  
 2021.07 ~ Now

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Award ID: 1911726  
 NSF-ANR ID: 10203871  
 Author(s) / Creator(s): Zhu, Bing-Yan, Hu, Zhong, Ying-Chen, Jinghui Du, Yanli, Cheng, Huihang, Yang, David A.  
 Date Published: 2020-02-01  
 Journal Name: Applied Mathematical Modelling  
 Volume: 78  
 Issue: C  
 ISSN: 0307-904X  
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Relativistic HPIC-LBM and its application in large temporal-spatial turbulent magnetic reconnection. Part II. Role of turbulence in the flux rope interaction

Award ID: 1911726  
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 Author(s) / Creator(s): Zhu, Bing-Yan, Hu, Zhong, Ying-Chen, Jinghui Du, Yanli, Cheng, Huihang, Yang, David  
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MLA

NPL&RAL  
 UK Space Weather and Space Environment Monitoring and Forecasting Programme  
 12-15 September 2023  
 Leonardo Hotel Cardiff, Cardiff, UK

欧洲空间天气会议  
 Software development of GeV-level-SEPs-induced extreme space weather disasters with plasma statistical physics theoretical model on domestic DCU accelerator heterogeneous supercomputer

20-24 NOVEMBER  
 EUROPEAN SPACE WEATHER WEEK  
 BRINGING SPACE WEATHER SPACE CLIMATE AND ENGINEERING TOGETHER

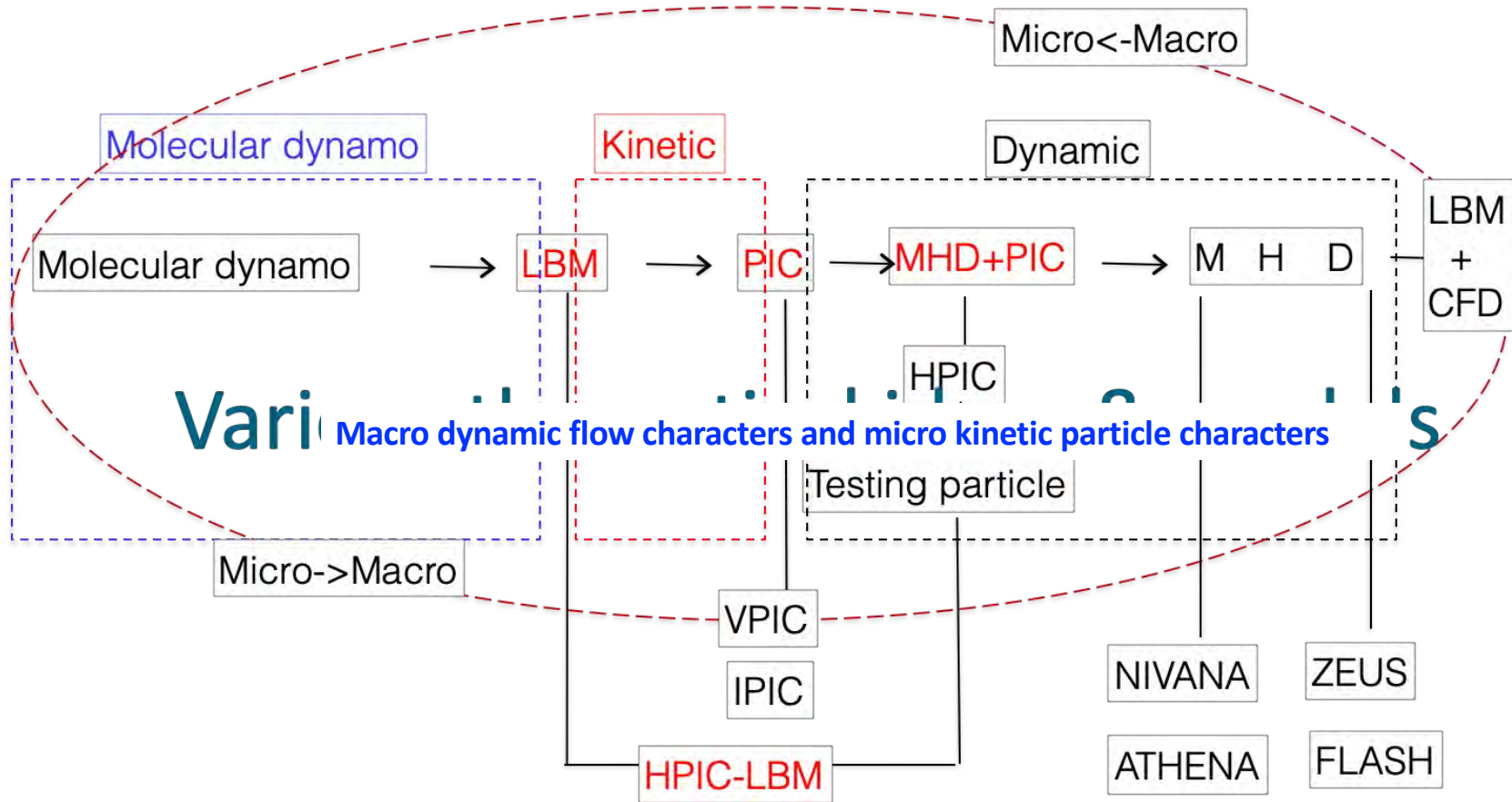
**Independent Intellectual Property Rights(IIPR)**

1983 Sauro Succi, 2009 Andrew Declan Diver, 2015 Mohseni Farhang (Hans J Herrminer, Sauro succi,Bastien Chopard) , 2017 Wolf-Gladrow Dieter(Neubauer F.M, Lussem M) et.al, Zhu et al 2007,2013, 2019,2020a,2020b,2023a,2023b,2023c

# PHPIC-LBM: Method

Non-ideal magnetic dissipation zone -MR

Lab    Earth-sun space    **Solar**    Interstellar medium (ISM)    Clusters of Galaxies



# PHPIC-LBM: Method

$$\frac{\partial f_{i\xi\ell}^{se}}{\partial t} + \frac{1}{\varepsilon_l^e} \mathbf{U}_l^e \cdot \nabla_x f_{i\xi\ell}^{se} + \mathbf{F}_l^e \cdot \frac{\partial f_{i\xi\ell}^{se}}{\partial \mathbf{p}_l^e} = \frac{\theta_l^e}{(\varepsilon_l^e)^2} \nabla_{\mathbf{p}_l^e} \cdot \left( \nabla_{\mathbf{p}_l^e} f_{i\xi\ell}^{se} + \mathbf{U}_l^e f_{i\xi\ell}^{se} \right) \quad (8)$$

$$\frac{\partial f_{i\xi\ell}^{si}}{\partial t} + \frac{1}{\varepsilon_l^i} \mathbf{U}_l^i \cdot \nabla_x f_{i\xi\ell}^{si} + \mathbf{F}_l^i \cdot \frac{\partial f_{i\xi\ell}^{si}}{\partial \mathbf{p}_l^i} = \frac{\theta_l^i}{(\varepsilon_l^i)^2} \nabla_{\mathbf{p}_l^i} \cdot \left( \nabla_{\mathbf{p}_l^i} f_{i\xi\ell}^{si} + \mathbf{U}_l^i f_{i\xi\ell}^{si} \right) \quad (9)$$

$$\frac{\partial f_{i\xi\ell}^{sn}}{\partial t} + \mathbf{U}_l^n \cdot \nabla_x f_{i\xi\ell}^{sn} + \mathbf{F}_l^n \cdot \frac{\partial f_{i\xi\ell}^{sn}}{\partial \mathbf{p}_l^n} = \nabla_{\mathbf{p}_l^n} \cdot \left( \nabla_{\mathbf{p}_l^n} f_{i\xi\ell}^{sn} + \mathbf{U}_l^n f_{i\xi\ell}^{sn} \right) \quad (10)$$

$$\frac{\partial \bar{\mathbf{B}}}{\partial t} = \nabla \times (\bar{\mathbf{U}} \times \bar{\mathbf{B}}) - \nabla \times (\hat{\alpha} \bar{\mathbf{B}}) + \hat{\beta} \nabla^2 \bar{\mathbf{B}} + \mu^{-1} (\eta_s + \eta_{MT} + \eta_{FT}) \nabla^2 \bar{\mathbf{B}} \quad (11)$$

$$\frac{\partial \tilde{\mathbf{B}}}{\partial t} = \nabla \times (\bar{\mathbf{U}} \times \tilde{\mathbf{B}}) + \nabla \times (\tilde{\mathbf{U}} \times \bar{\mathbf{B}}) + \nabla \times (\tilde{\mathbf{U}} \times \tilde{\mathbf{B}}) - \nabla \times (\hat{\alpha} \bar{\mathbf{B}}) + \hat{\beta} \nabla^2 \bar{\mathbf{B}} + \mu^{-1} (\eta_s + \eta_{MT} + \eta_{FT}) \nabla^2 \tilde{\mathbf{B}} \quad (12)$$

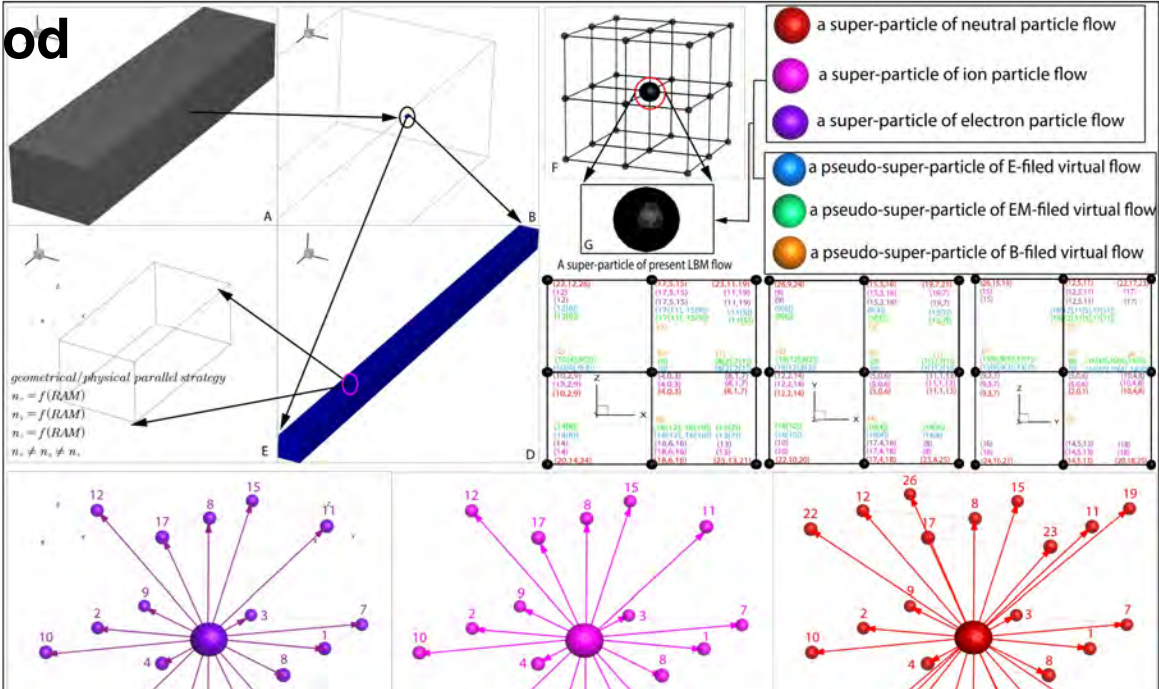
$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times \left\{ \begin{array}{l} \frac{m_e}{e} \frac{\partial \mathbf{U}_e}{\partial t} + \frac{m_e}{e} (\mathbf{U}_e \nabla) \mathbf{U}_e + \frac{1}{en_e} \nabla P_e + \frac{1}{en_e} \nabla \pi_e + \frac{m_e}{e} \langle \xi \cdot \nabla \cdot \xi \rangle + \mathbf{U}_e \times \mathbf{B} - \alpha_{\parallel} (\mathbf{J} \cdot \mathbf{b}) \mathbf{b} - \\ \alpha_{\perp} \mathbf{b} \times (\mathbf{J} \times \mathbf{b}) + \alpha_{\Lambda} (\mathbf{b} \times \mathbf{J}) + \frac{\beta_{\parallel}}{e} (\nabla T_e \cdot \mathbf{b}) \mathbf{b} + \frac{\beta_{\perp}}{e} \mathbf{b} \times (\nabla T_e \times \mathbf{b}) + \frac{\beta_{\Delta}}{e} (\mathbf{b} \times \nabla T_e) + \langle \xi \times \tilde{\mathbf{B}} \rangle \end{array} \right\} \quad (13)$$

$$\frac{\partial f_{1\xi\ell}^{sn}}{\partial t} + \frac{1}{\varepsilon_1^n} \mathbf{U}_1^n \cdot \nabla_x f_{1\xi\ell}^{sn} + \mathbf{F}_1^n \cdot \frac{\partial f_{1\xi\ell}^{sn}}{\partial \mathbf{p}_1^n} = \frac{\theta_1^n}{(\varepsilon_1^n)^2} \nabla_{\mathbf{p}_1^n} \cdot \left( \nabla_{\mathbf{p}_1^n} f_{1\xi\ell}^{sn} + \mathbf{U}_1^n f_{1\xi\ell}^{sn} \right) \quad (15)$$

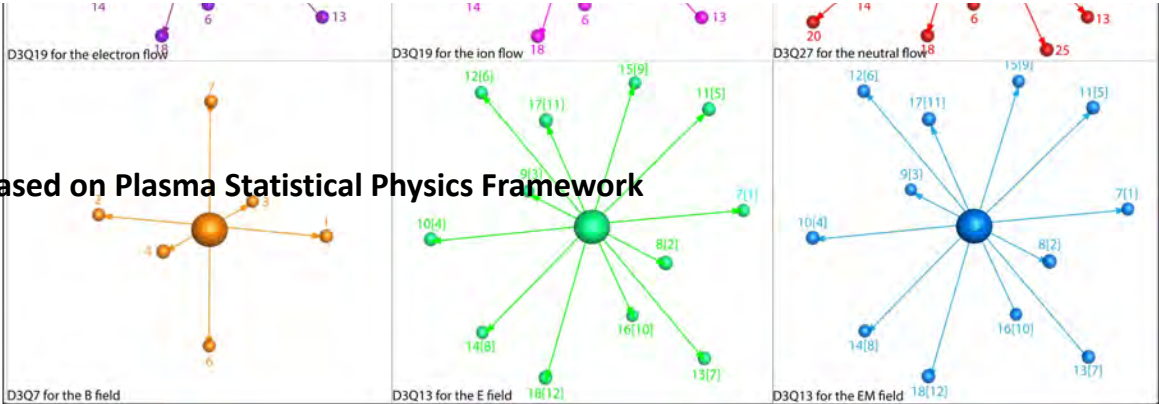
$$\mathbf{F}_1^n = \nabla \cdot [\rho \mathbf{U}_1 \mathbf{U}_1 \mathbf{I} + p \mathbf{I} - \xi_n^{TVD} \nabla \cdot \mathbf{U}_1 - \rho g] - \nabla \cdot \left\{ (\vartheta_{n0}^{VD} + \xi_n^{TVD}) \left[ \nabla \mathbf{U}_1 + (\nabla \mathbf{U}_1)^{\top} - \frac{2 \nabla \cdot \mathbf{U}_1}{3} \mathbf{I} \right] \right\} \quad (16)$$

The Relativistic HPIC-LBM Code based on Plasma Statistical Physics Framework

# PHPIC-LBM: Method



Convert the interaction between fields and real flow into pseudo-flow(fields) and real flow through lattice grid model



## The Relativistic HPIC-LBM Code based on Plasma Statistical Physics Framework

# RHPIC-LBM Method Turbulence induced viscosity & resistivity dissipation and diffusion model

**Definition:** eddy-viscosity type resistivity coefficient MHD magnetic dissipation region of LSTMR

$$\alpha = -\frac{1}{3} \int_{-\infty}^t \langle \tilde{V} \cdot \nabla \times \tilde{V}' \rangle dt' = -\frac{1}{3} \tau_\alpha \langle \tilde{V} \cdot \nabla \times \tilde{V} \rangle = -\frac{1}{3} \tau_\alpha H^K = -\frac{1}{3} \tau_\alpha \int_{Volume} (\mathbf{V} \cdot \nabla \times \mathbf{V}) dx^3$$

$\mathbf{V} \cdot \nabla \times \mathbf{V}$  The helicity density  $H^K$  The kinetic helicity

The mean component of velocity and magnetic field parts-----pseudo scaler-----B

$$\eta_t^M = \frac{1}{3} \int_{-\infty}^t \langle \tilde{V} \cdot \tilde{V}' \rangle dt' = \frac{1}{3} \tau_\beta \langle \tilde{V}^2 \rangle = \frac{1}{3} \tau_\beta E^K$$

$E^K$  The kinetic energy of turbulence

The fluctuating component of velocity and magnetic field parts-----pseudo scaler-----J

$$\eta_t^V \sim \omega \quad \omega = \nabla \times \mathbf{V}$$

The fluctuating component of velocity and magnetic field parts-----pseudo scaler----- $\omega$

**Definition:** Spitzer resistivity

$$\eta_s = \frac{\pi Z e^2 \ln \Lambda \sqrt{m_e}}{(4\pi\epsilon_0)^2 (k_B T_e)^{3/2}} = \frac{\eta_c^2}{4\pi}$$

$$|\eta_s| = 1.96 \eta_{s0} = 1.03 \times 10^{-4} T_e^{-3/2} Z \ln \Lambda \approx 1.03 \times 10^{-4} \ln \left( \frac{Z^{1.71828} T_e^{1.72313}}{m_e} \right)$$

Viscous induced dissipation (VD,  $\vartheta_{n0}$ ,  $\xi_n^{PT}$ ,  $\xi_n^{TT}$ ,  $\vartheta_{i0}$ ,  $\xi_i^{PT}$ ,  $\xi_i^{TT}$ ,  $\vartheta_{e0}$ ,  $\xi_e^{PT}$ ,  $\xi_e^{TT}$ , see Appendix B and

1<sup>st</sup> type: Plasma bulk kinetic energy ( $E_U$ ) dissipation into plasma heating energy ( $E_T$ ) and non-thermal kinetic energy that accelerate particles ( $E_{SEP}$ , here we will focus on the solar energetic particles, SEP)(e.g., Duijveman 1981; Ryan et al. 2000; Tandberg-Hanssen 2009; Antolin et al. 2020; Pontinet al. 2020);

2<sup>nd</sup> type:  $E_U$  dissipation into magnetic energy ( $E_B$ ).

## Resistivity and Viscosity

**Definition:** non-ideal MHD magnetic dissipation region of LSTMR

Mathematical model: assuming that magnetic dissipation effects depend mainly on the **current density** a Taylor expansion gives

$$\mathbf{E} + \mathbf{V} \times \mathbf{B} = \eta_1 \mathbf{j} - \eta_2 \nabla^2 \mathbf{j} + \dots \quad \eta_1 = \left( \frac{c}{\omega_{pe}} \right)^2 \nu_e \quad \eta_2 \simeq \left( \frac{c}{\omega_{pe}} \right)^2 \mu_e$$

$\nu_e$  Electron-ion collision frequency  $\mu_e \sim r_e^2 \nu_e$  Electron viscosity

$$\omega_{pe} = e \sqrt{4\pi n / m_e} \quad \text{Electron plasma frequency}$$

Physical model: Dividing the fields into mean and fluctuating parts, the mean turbulence energy-dissipation rate closed linked eddy-viscosity can be written as

$$\varepsilon = \alpha \langle \mathbf{B} \rangle \cdot \left[ \eta_t^M \langle \mathbf{j} \rangle + \eta_t^V \langle \boldsymbol{\omega} \rangle \right] + \alpha_t \langle \mathbf{B} \rangle \cdot \eta_t^M \langle \nabla \times \mathbf{B} \rangle + \eta_t^V \langle \nabla \times \mathbf{V} \rangle$$

Physical model: split the velocity and magnetic field into mean and fluctuating parts, from mean field kinematic dynamo theory, the induction equation can be written as

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\mathbf{V} \times \mathbf{B}) + \nabla \times (\tilde{\mathbf{V}} \times \tilde{\mathbf{B}}) + \eta_s \nabla^2 \tilde{\mathbf{B}}$$

Resistive induced dissipation (RD,  $\eta_n^V$ ,  $\eta_n^t$ ,  $\eta_i^t$ ,  $\eta_e^t$ , see Appendix B and following sections 2.1 for def-

According to the classification criteria of resistivity properties of solar plasma, the energy dissipation in this process can be categorized into two different categories depending on the coherent structures of  $\mathbf{B}$ -turbulence:

3<sup>rd</sup> type:  $E_B$  dissipation into  $E_T$  and  $E_{SEP}$ ;

4<sup>th</sup> type:  $E_B$  dissipation into  $E_U$ .

# RHPIC-LBM Method Turbulence induced viscosity & resistivity dissipation and diffusion model

CLS	$L_{LE}$	$L_{UA}^M$	$L_{\xi_n}^{PT}$	$L_{\xi_n}^{TT}$	$L_{\eta_n}^S$	$L_{\eta_n}^T$	$L_i$	$L_{\xi_i}^{PT}$	$L_{\xi_i}^{TT}$	$L_{\eta_i}^T$	$L_{UA}^K$	$L_e$	$L_{\xi_e}^{PT}$	$L_{\xi_e}^{TT}$	$L_{\eta_e}^T$
<b>I(1-3): (<math>l &gt; L_{\xi_n}^{PT}</math>), FC, LEM, KLS</b>															
<b>I(1): Electric neutral fluid (<math>l &gt; L_{LE} \gg L_{UA}^M</math>)</b>															
<b>I(2): Ideal electrically conducting MHD (<math>L_{LE} &gt; l &gt; L_{UA}^M</math>),</b>															
<b>I(3): Quasi-neutral ideal MHD (<math>L_{UA}^M &gt; l &gt; L_{\xi_n}^{PT}</math>)</b>															
$EC_{dyn}  _{HD}$	$(\checkmark^a)$	$I(1) \times \checkmark, I(2) \checkmark$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
$EC_{dyn}  _{MHD}$	$\times$	$I(3) \checkmark$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
<b>II(1-2): (<math>L_{\xi_n}^{PT} &gt; l &gt; L_{\eta_n}^S</math>), FC</b>															
<b>II(1): Quasi-neutral non-ideal MHD (<math>L_{\xi_n}^{PT} &gt; l &gt; L_{\xi_n}^{TT}</math>), KLS, VD/TVD [<math>\xi_n^{PT}</math>]</b>															
<b>II(2): WIP non-ideal MHD (<math>L_{\xi_n}^{TT} &gt; l &gt; L_{\eta_n}^S</math>), TAC, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}</math>]</b>															
$VD_\theta  _{HD}$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
$VD_\xi  _{HD}$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
<b>III(1-3): (<math>L_{\eta_n}^S &gt; l &gt; L_{\xi_i}^{PT}</math>), TAC</b>															
<b>III(1): WIP non-ideal MHD to PIP non-ideal MHD (<math>L_{\eta_n}^S &gt; l &gt; L_{\eta_n}^T</math>), FC, VD/TVD [<math>\xi_n^{PT}</math>], RD/TRD [<math>\eta_n^S</math>]</b>															
<b>III(2): PIP non-ideal MHD (<math>L_{\eta_n}^T &gt; l &gt; L_i</math>), FC&amp;PC, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T</math>]</b>															
<b>III(3): FIP MHD (<math>L_i &gt; l &gt; L_{\xi_i}^{PT}</math>), FC&amp;PC, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T</math>], HMHD [<math>L_i</math>]</b>															
$VD_\xi  _{MHD}$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
$RD_{\eta_n}  _{MHD}$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
$RD_{\eta_n^T}  _{MHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
$HMDD$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$
<b>IV(1-2): (<math>L_{\xi_i}^{PT} &gt; l &gt; L_{\eta_i}^T</math>), FIP MHD, FC&amp;PC, TAC</b>															
<b>IV(1): (<math>L_{\xi_i}^{PT} &gt; l &gt; L_{\xi_i}^{TT}</math>), VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T</math>]</b>															
<b>IV(2): (<math>L_{\xi_i}^{TT} &gt; l &gt; L_{\eta_i}^T</math>), VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T, \eta_i^T</math>]</b>															
$VD_\theta  _{IMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$
$VD_\xi  _{IMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$
$RD_{\eta_n^T}  _{IMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$
<b>V(1-3): (<math>L_{\eta_n}^T &gt; l &gt; L_{\xi_e}^{PT}</math>), FIP MHD, FC&amp;PC, TAC</b>															
<b>V(1): (<math>L_{\eta_n}^T &gt; l &gt; L_{UA}^K</math>), FIP MHD, FC&amp;PC, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T, \eta_i^T</math>], TAC</b>															
<b>V(2): (<math>L_{UA}^K &gt; l &gt; L_e</math>) FIP MHD, FC&amp;PC, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T, \eta_i^T</math>], TAC</b>															
<b>V(3): (<math>L_e &gt; l &gt; L_{\xi_e}^{PT}</math>) EMHD, FC&amp;PC, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T, \eta_i^T</math>], TAC, MF</b>															
$VD_\xi  _{IMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$
$RD_{\eta_n^T}  _{IMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$	$\times$
<b>VI(1-2): (<math>L_{\xi_e}^{PT} &gt; l &gt; L_{\eta_e}^T</math>), FIP MHD, FC&amp;PC, TAC</b>															
<b>VI(1): (<math>L_{\xi_e}^{PT} &gt; l &gt; L_{\xi_e}^{TT}</math>) EMHD, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}, \xi_e^{PT}, \xi_e^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T, \eta_i^T</math>]</b>															
<b>VI(2): (<math>L_{\xi_e}^{TT} &gt; l &gt; L_{\eta_e}^T</math>) EMHD, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}, \xi_e^{PT}, \xi_e^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T, \eta_i^T</math>]</b>															
$VD_\theta  _{EMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$
$VD_\xi  _{EMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$
$RD_{\eta_n^T}  _{EMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$	$\checkmark$	$\checkmark$	$\times$	$\times$	$\times$
<b>VII (<math>L_{\eta_n}^T &gt; l</math>) EMHD, FC&amp;PC, VD/TVD [<math>\xi_n^{PT}, \xi_n^{TT}, \xi_i^{PT}, \xi_i^{TT}, \xi_e^{PT}, \xi_e^{TT}</math>], RD/TRD [<math>\eta_n^S, \eta_n^T, \eta_i^T, \eta_e^T</math>], TAC</b>															
$RD_{\eta_n^T}  _{EMHD}$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\checkmark$

<sup>a</sup>  $\checkmark$ : Energy dissipation (e.g., EC, VD/TVD, RD/TRD, Hall effect dissipation, and so on) happen at this special CLS.

<sup>b</sup>  $\times$ : Energy dissipation (e.g., EC, VD/TVD, RD/TRD, Hall effect dissipation, and so on) don't happen at this special CLS.

NOTE—For symbol definitions, see Appendix B.

Table 4. GR and PPC under different scales

$\varrho$	$\xi$	Hurst index, $H_\varrho$	GR (cm)	PPC <sup>a</sup>	Space Unit (cm)	Time Unit (s)	Frequency Unit (rad.s <sup>-1</sup> )
1	1	$H_1$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$\varrho^{\tilde{a}(H_1)}$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$1000\omega_{pi}^{-1} \sim 3.2 \times 10^{-3}$	$\times^b$
2 ~ 3	3	$H_2^c, H_3^d$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$\varrho^{\tilde{a}(H_2)^c, \tilde{a}(H_3)^d}$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$\omega_{pi}^{-1} \sim 3.2 \times 10^{-7}$	$\omega_{pi} \sim 3.1 \times 10^6$
4	2	$H_4$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$\varrho^{\tilde{a}(H_4)}$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$\omega_{pi}^{-1} \sim 3.2 \times 10^{-7}$	$\omega_{pi} \sim 3.1 \times 10^6$
5	2	$H_5$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$\varrho^{\tilde{a}(H_5)}$	$c\omega_{pi}^{-1} \sim 9 \times 10^3$	$\omega_{pi}^{-1} \sim 3.2 \times 10^{-7}$	$\omega_{pi} \sim 3.1 \times 10^6$
6	2	$H_6$	$c\omega_{pe}^{-1} \sim 5^e$	$\varrho^{\tilde{a}(H_6)}$	$c\omega_{pe}^{-1} \sim 5$	$\omega_{pe}^{-1} \sim 0.17 \times 10^{-9}$	$\omega_{pe} \sim 5.64 \times 10^9$

<sup>a</sup> The Hurst index determined the PPC.

<sup>b</sup>  $\times$ : No frequency Unit at this scale.

<sup>c</sup> For neutrals fluid flow.

<sup>d</sup> For electron fluid flow and ion fluid flow.

<sup>e</sup> Same order with electron Debye Length inside the current sheet,  $\lambda_{De}$  (e.g., Narita et al. 2020; Maiorano et al. 2020; Settiou 2020).

NOTE—For symbol definitions, see Appendix B.

Table 5. Non-dimensional viscosity and resistivity under different scales

$\varrho$	$\xi$	$\vartheta_{n0}$	$\vartheta_n^{PT}$	$\xi_n^{TT}$	$\vartheta_{i0}$	$\vartheta_i^{PT}$	$\xi_i^{TT}$	$\vartheta_{e0}$	$\vartheta_e^{PT}$	$\vartheta_e^{TT}$	$\eta_n^S$	$\eta_n^T$	$\eta_i^T$	$\eta_e^T$
1	1	$\times 10^{-4}$	$\frac{C_1 E_{\theta n}^{PT}}{\epsilon_{\theta n}^2}$	$\frac{C_2 E_{\theta n}^{TT}}{\epsilon_{\theta n}^2}$	$\times^a$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$	$\times$
2, 3	3	$\times 10^{-4}$	$\frac{C_1 E_{\theta n}^{PT}}{\epsilon_{\theta n}^2}$	$\frac{C_2 E_{\theta n}^{TT}}{\epsilon_{\theta n}^2}$	$\frac{m_i(\kappa T_i)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	$\frac{C_1 E_{\theta i}^{PT}}{\epsilon_{\theta i}^2}$	$\frac{C_2 E_{\theta i}^{TT}}{\epsilon_{\theta i}^2}$	$\frac{m_e(\kappa T_e)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	$\frac{C_2 E_{\theta e}^{PT}}{\epsilon_{\theta e}^2}$	$\frac{C_2 E_{\theta e}^{TT}}{\epsilon_{\theta e}^2}$	$\eta_n^S b, \eta_n^S c$	$\frac{1}{4\pi\sigma_T} d$	$\frac{1}{4\pi\sigma_T} e$	$\frac{1}{4\pi\sigma_T} f$
4	2	$\times$	$\times$	$\times$	$\frac{m_i(\kappa T_i)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	1	1	$\frac{m_e(\kappa T_e)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	1	1	$\eta_n^S, \eta_n^S c$	$\times$	$\frac{1}{4\pi\sigma_T}$	$\frac{1}{4\pi\sigma_T}$
5	2	$\times$	$\times$	$\times$	$\frac{m_i(\kappa T_i)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	1	1	$\frac{m_e(\kappa T_e)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	1	1	$\eta_n^S, \eta_n^S c$	$\times$	$\frac{1}{4\pi\sigma_T}$	$\frac{1}{4\pi\sigma_T}$
6	2	$\times$	$\times$	$\times$	$\frac{m_i(\kappa T_i)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	1	1	$\frac{m_e(\kappa T_e)^{\frac{3}{2}}}{C_3 e^4 l n \Lambda}$	1	1	$\eta_n^S, \eta_n^S c$	$\times$	$\frac{1}{4\pi\sigma_T}$	$\frac{1}{4\pi\sigma_T}$

<sup>a</sup> The PPC is determined by the Hurst index.

<sup>b</sup>  $\eta_{ni}^S = \eta_0 + \eta_1 \frac{|J \times B| |U_e|^2}{|B|^2}$ ,  $\eta_0 = 5.0 \times 10^{-4}$ ,  $\eta_1 = 1.0 \times 10^{-3}$  (e.g., Piddington 1981; Barghouty et al. 1996; Brandenburg et al. 2008; Hazra et al. 2018; Dikpati et al. 2007; Mark et al. 2014; Inoue et al. 2014; ?; Bamba et al. 2020; Miyoshi et al. 2020; Kusano et al. 2020).

<sup>c</sup>  $\eta_{ne}^S = \eta_0 + \eta_1 \frac{|J \times B| |U_e|^2}{|B|^2}$ .

<sup>d</sup>  $\sigma_T = \frac{\sqrt{3}\sigma_e}{\sqrt{3c^2 + 4\pi U_e L_\varrho}}$ ,  $U = \frac{U_n + U_i + U_e}{3}$

<sup>e</sup>  $\sigma_i^T = \frac{\sqrt{3}\sigma_e}{\sqrt{3c^2 + 4\pi U_i L_\varrho}}$

<sup>f</sup>  $\sigma_e^T = \frac{\sqrt{3}\sigma_e}{\sqrt{3c^2 + 4\pi U_e L_\varrho}}$

<sup>g</sup>  $l n \Lambda = l n \frac{2}{2e^2} \left\{ \begin{array}{l} \left( \frac{\kappa^3 T_e^2}{\pi n_e} \right)^{1/2} \quad T_e < 5.8 \times 10^5 \\ 100 \times \left( \frac{58 \kappa^3 T_e^2}{\pi n_e} \right)^{1/2} \quad T_e > 5.8 \times 10^5 \end{array} \right.$  (e.g., Chapman 1953; Leontovich 1967; Somov 1991, 2008, 2013; Gordovskyy 2020; Muhamad 2020)

NOTE—For symbol definitions, see Appendix B.

# RHPIC-LBM Method

Multi-component-abundance-isotope model

Table 1. Chemical composition of the Sun

Chemical composition	Percent of the overall mass	Type
<b>Hydrogen</b>	<b>73%</b>	<b>Dominated</b>
<b>Helium</b>	<b>25%</b>	<b>Dominated</b>
Oxygen	0.8%	Metals
Carbon	0.36%	Metals
Iron	0.16%	Metals
Neon	0.10%	Metals
Electron, Proton, $3\text{He}$ , $4\text{He}$ with real proton-to-electron mass ratio & neutron-electron mass ratio		
Hydrogen	0.02%	Metals
Silicon	0.07%	Metals
Magnesium	0.05%	Metals
Sulphur	0.04%	Metals
Others combined	0.04%	Metals

Table 2. Layers of the Sun

Layers	Thickness (km)	Temperature (K)	Type
Core	150,000	10,000,000	Inertial layer
Convection Zone	180,000	500,000	Inertial layer
Photosphere <sup>1</sup>	400	6400	Outer layer, un-magnetized, weakly ionized
Chromosphere	2100	4400 – 8000	Outer layer, magnetized, partly ionized <sup>2</sup>
Transition region	100	8000 – 500,000	Outer layer, magnetized, high ionized
Corona	5,000,000	1,000,000	Outer layer, magnetized, fully ionized

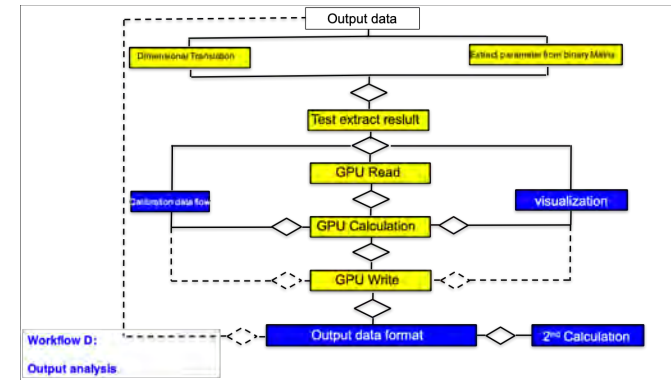
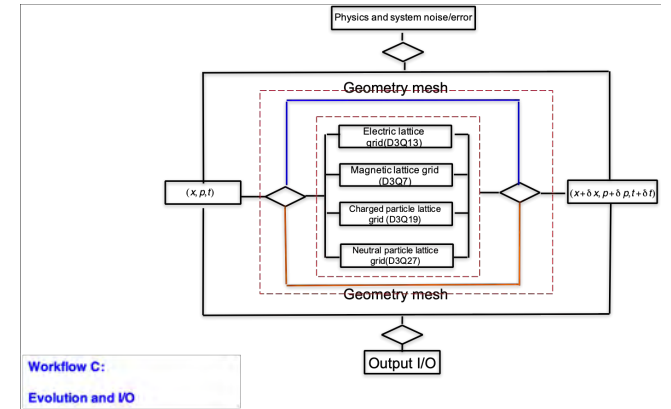
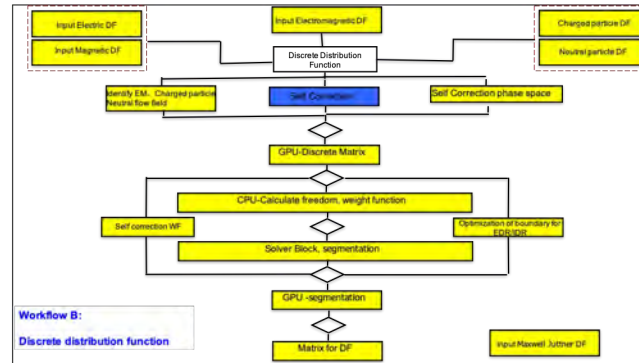
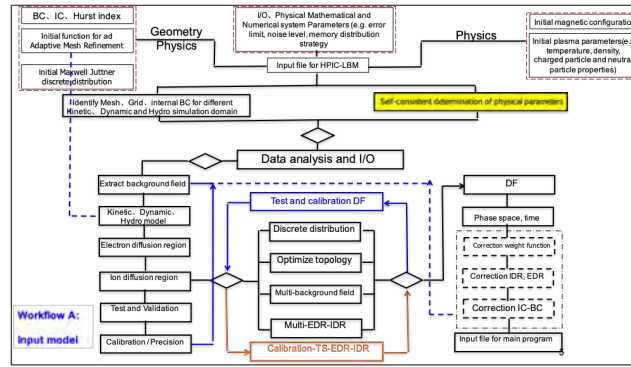
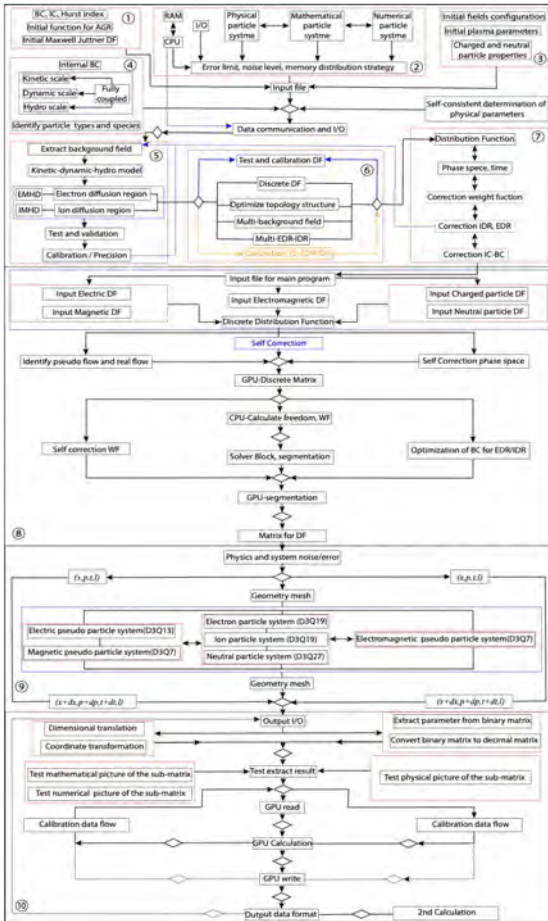
Multi-species & Multi-Degree of Ionization

# RHPIC-LBM Algorithm

- Few things we need to know before running codes

(OpenCL, 计算⑧、OpenGL, 可视化⑩)

## 程序框架图





# RHPIC-LBM code

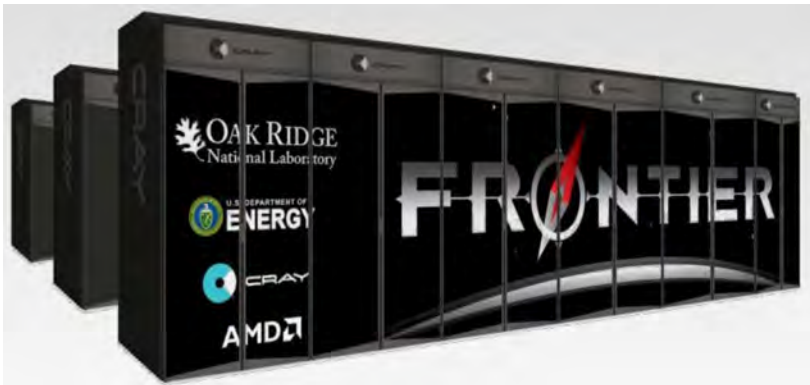
X86 CPU Version (2014~2018)



CPU-DCU (2019-2023.7)



CPU-GPU Version



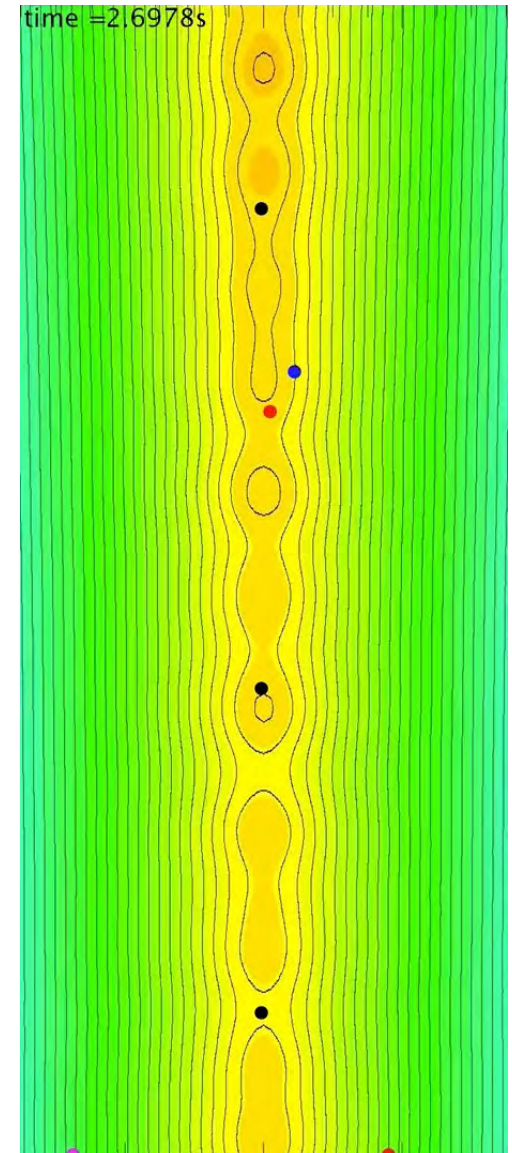
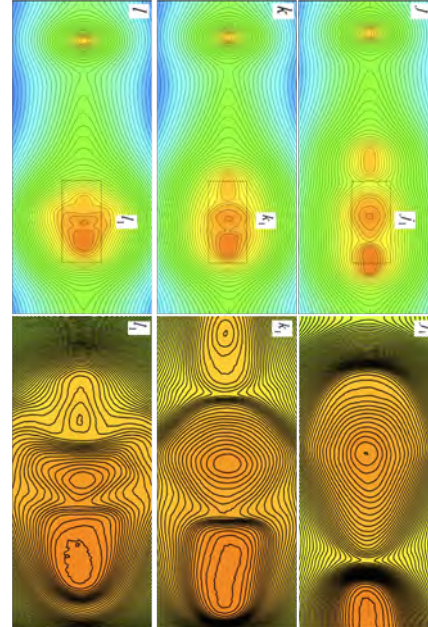
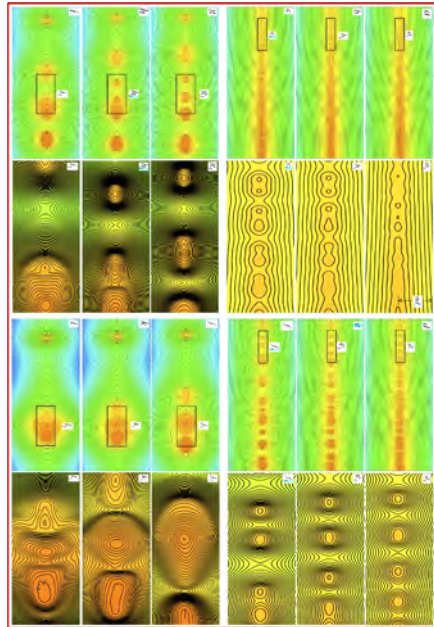
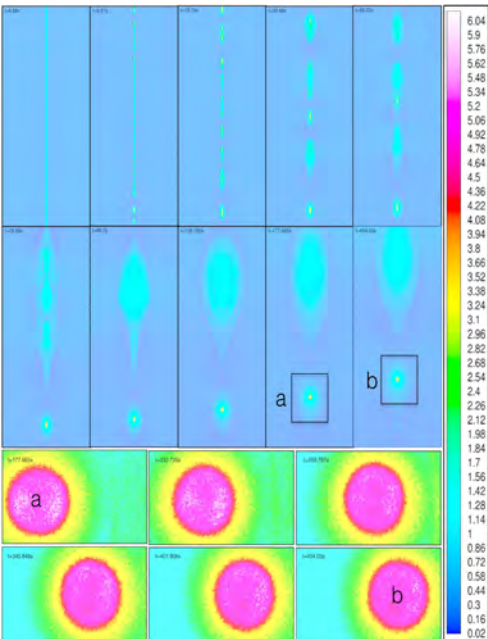
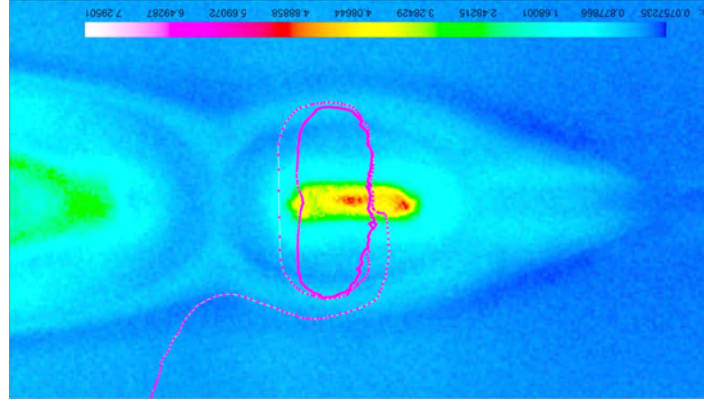
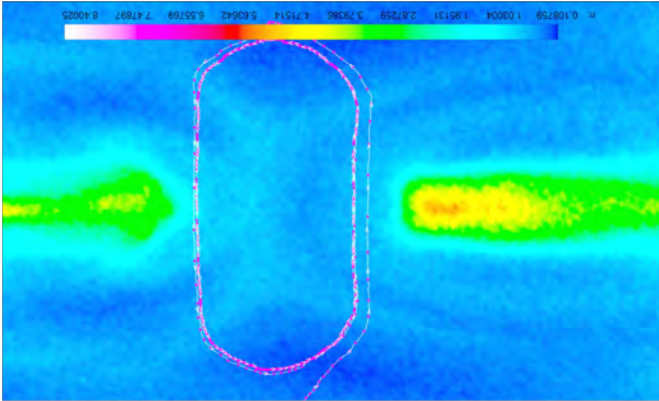
ARM CPU (2023~)



# RHPIC-LBM validation

Hybrid Acceleration Diffusive shock acceleration (DSA, 1st Fermi) & Stochastic acceleration (2nd Fermi)

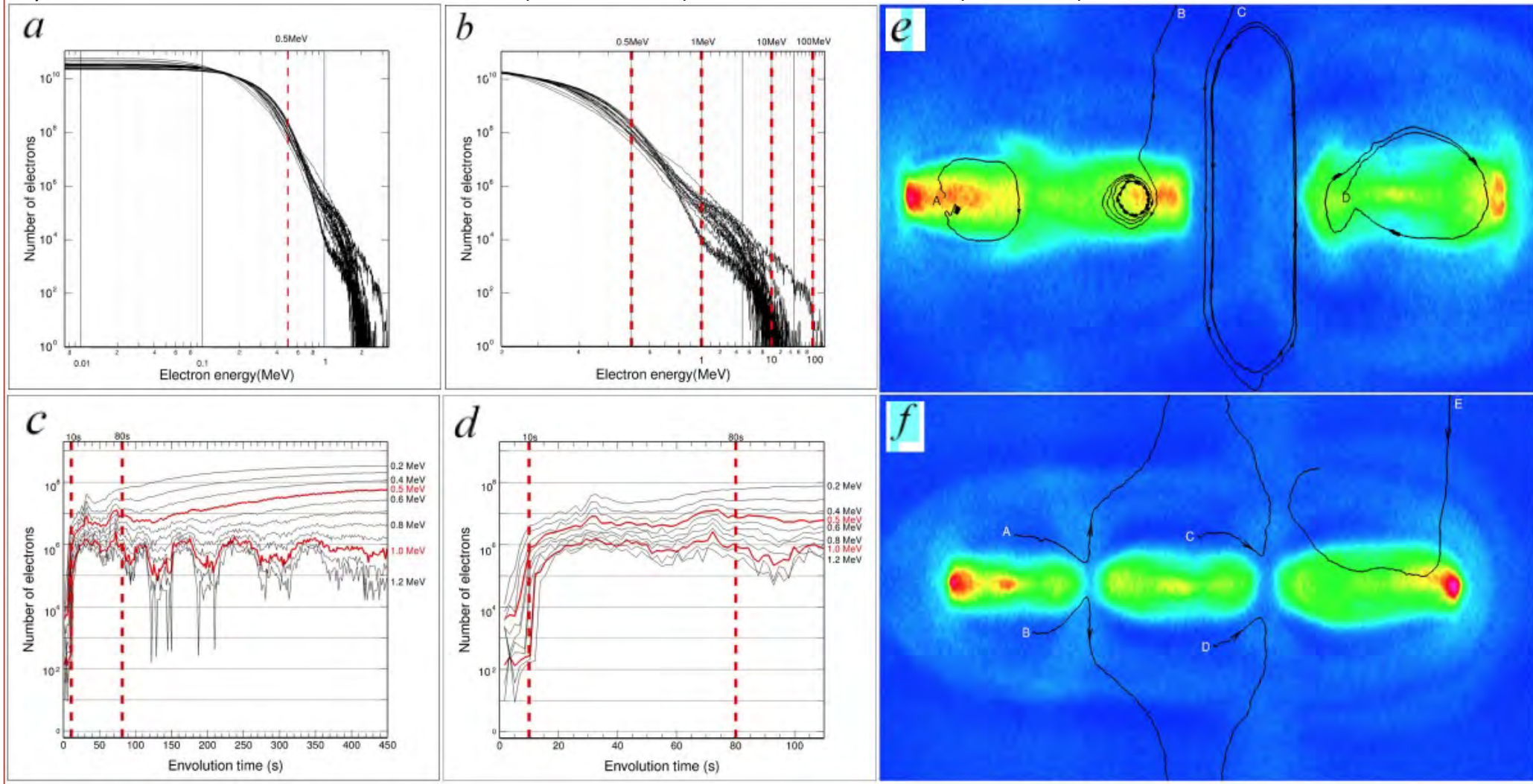
## 2.5D turbulence acceleration model I



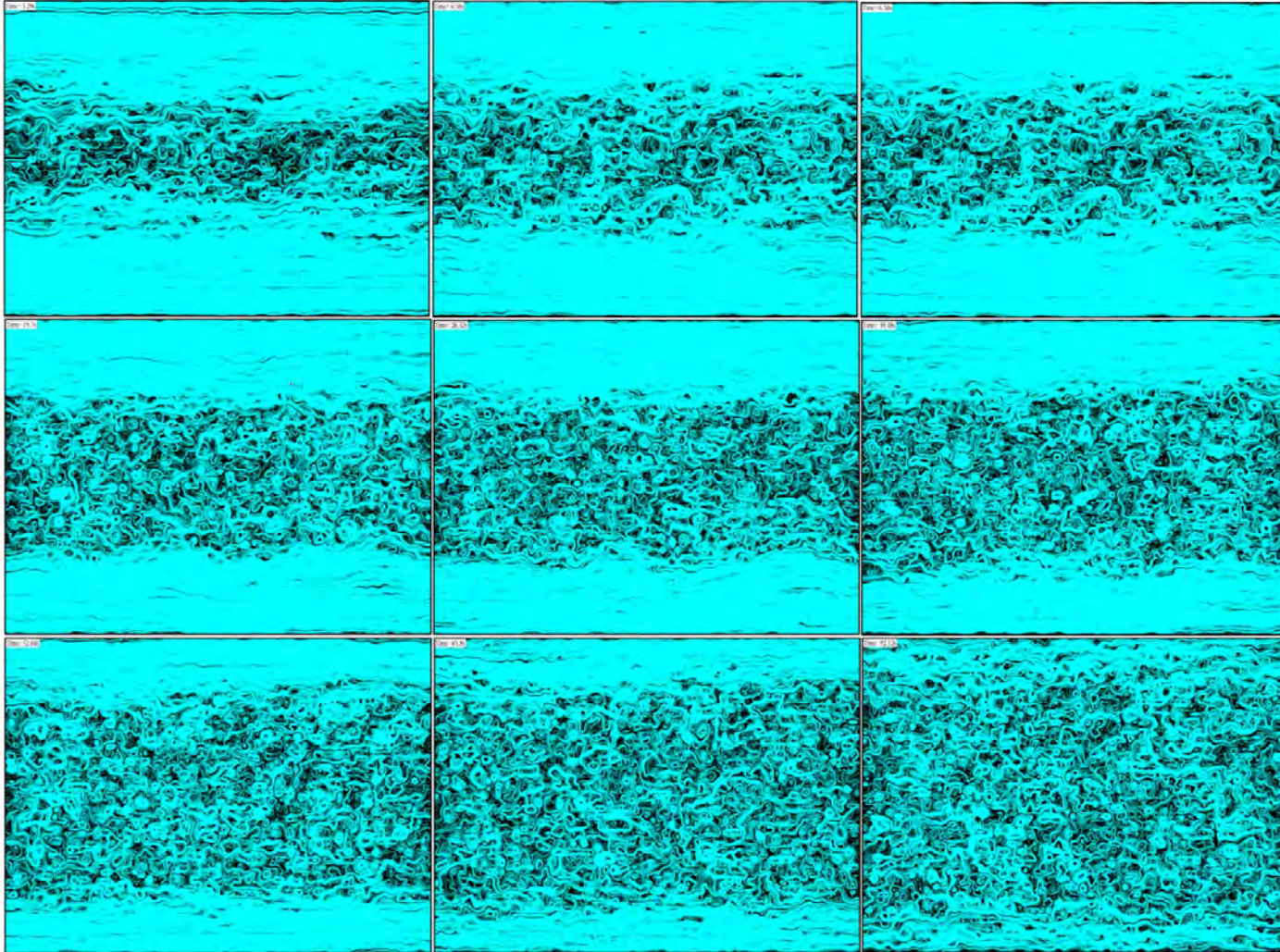
# RHPIC-LBM validation

## 2.5D turbulence acceleration model I

Hybrid Acceleration Diffusive shock acceleration (DSA, 1st Fermi) & Stochastic acceleration (2nd Fermi)



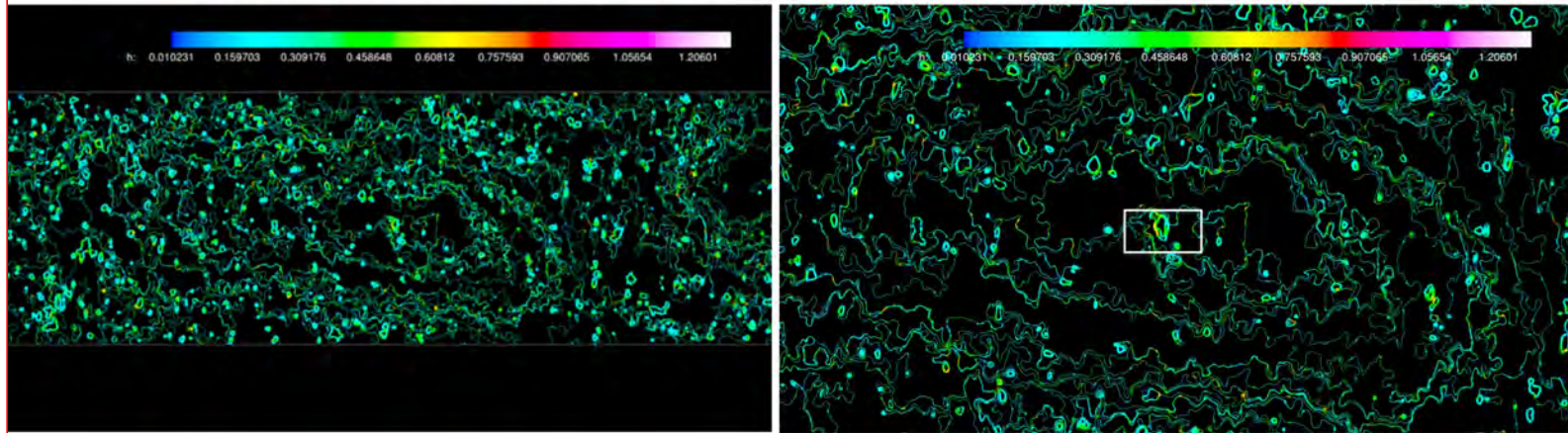
# RHPIC-LBM validation 2.5D turbulence acceleration model II-Fractal CS



URL: <https://pan.baidu.com/s/1ALXiDA4hEZhEES93p4o9nQ>

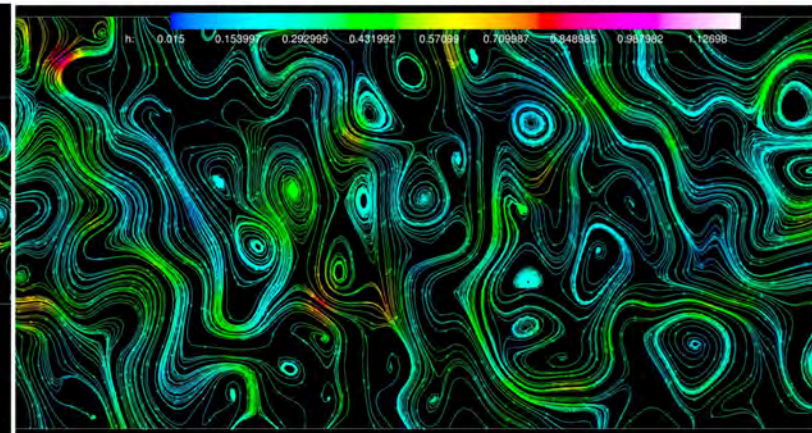
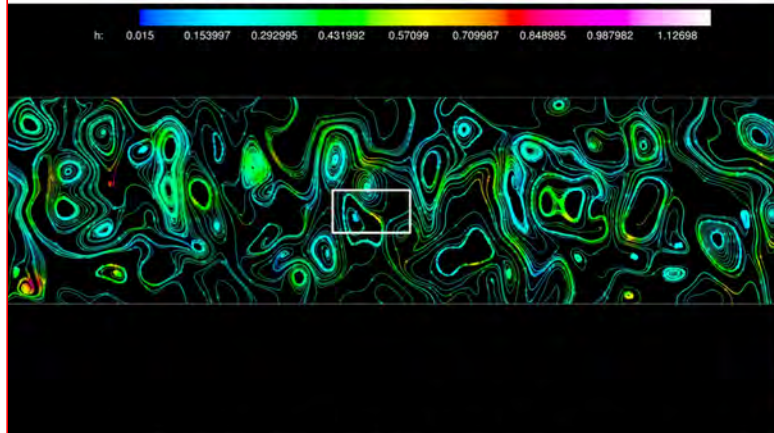
# RHPIC-LBM validation 2.5D turbulence acceleration model II-Fractal CS

Diffusive shock acceleration (DSA, 1st Fermi) & Stochastic acceleration (2nd Fermi)



Simulation Domain  $\sim 10^4 - 10^5 km$

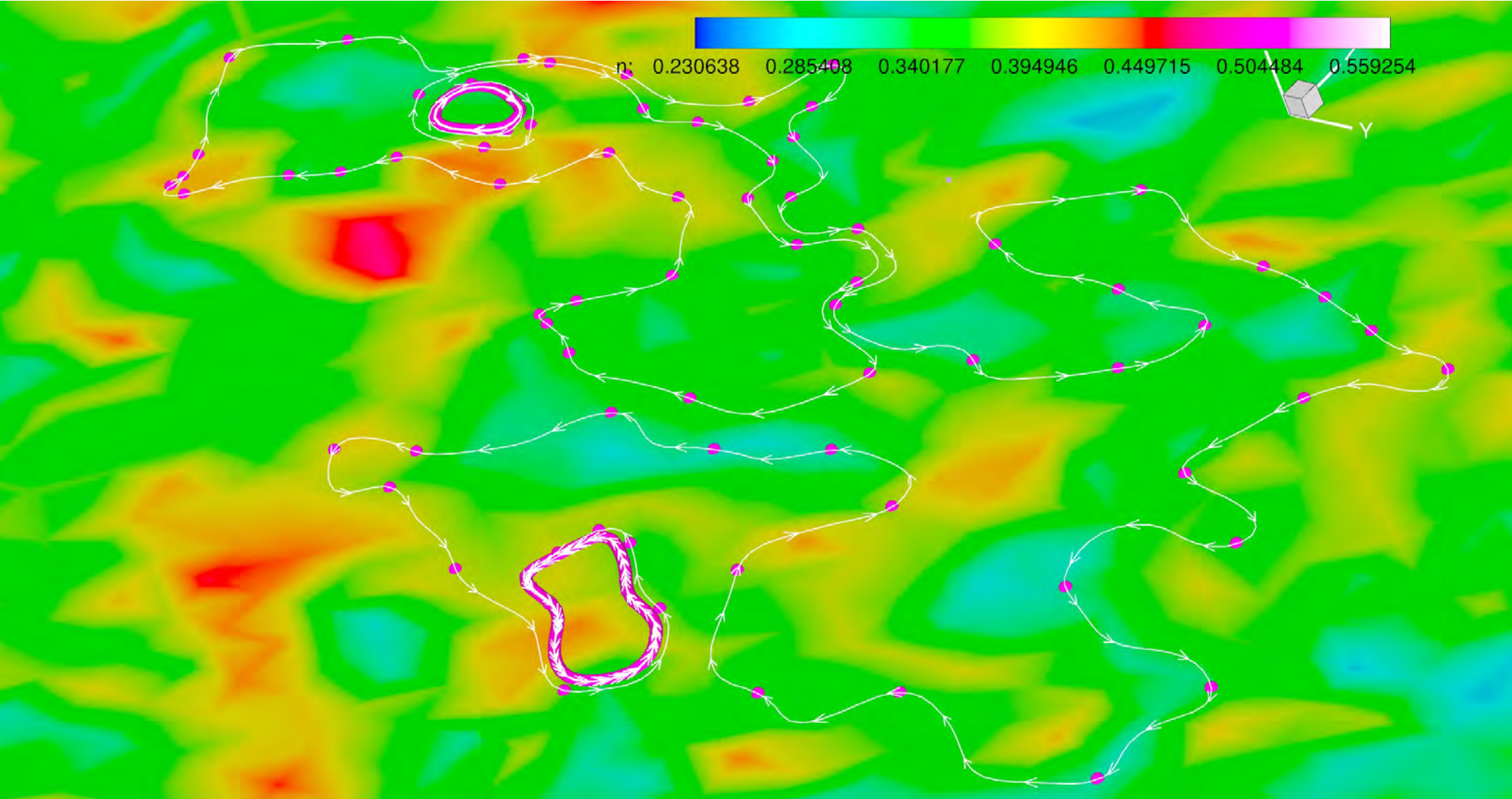
Simulation Domain  $\sim 10^3 - 10^4 km$



Simulation Domain  $\sim 10^2 - 10^3 km$

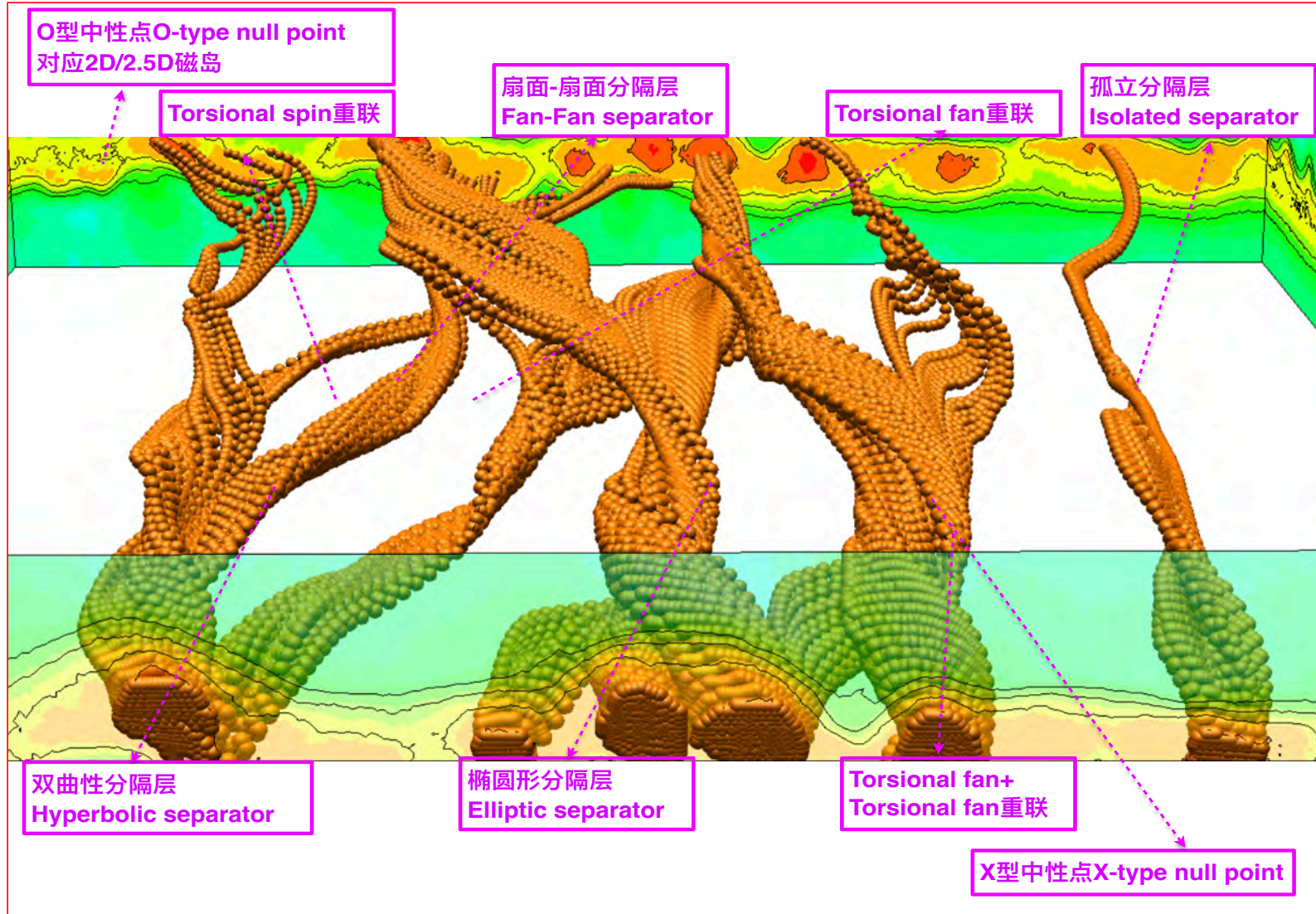
Simulation Domain  $\sim 10^1 - 10^2 km$

# RHPIC-LBM validation 2.5D turbulence acceleration model II-Fractal CS



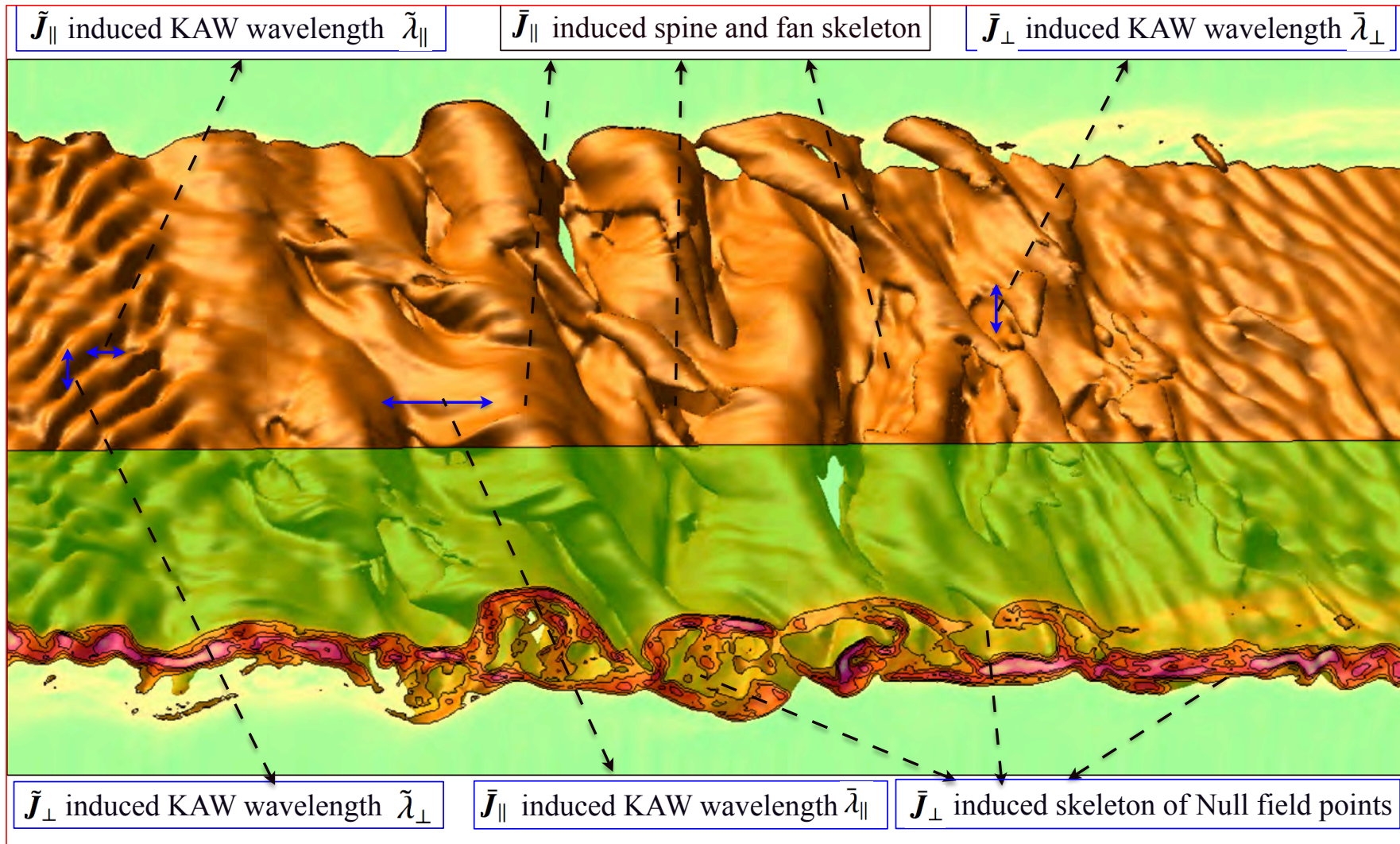
# RHPIC-LBM validation

## 3D turbulence acceleration model I



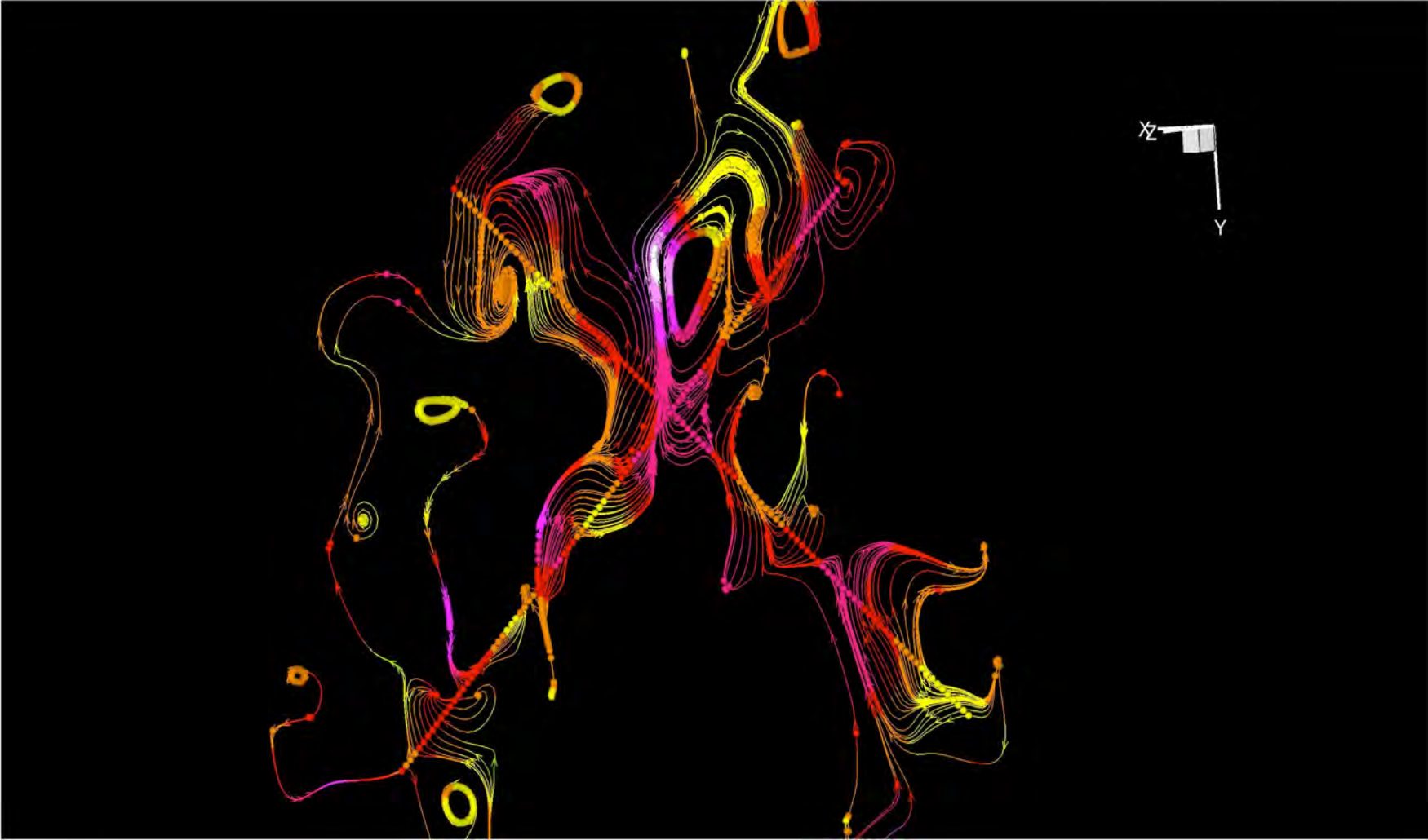
# RHPIC-LBM validation

## 3D turbulence acceleration model I



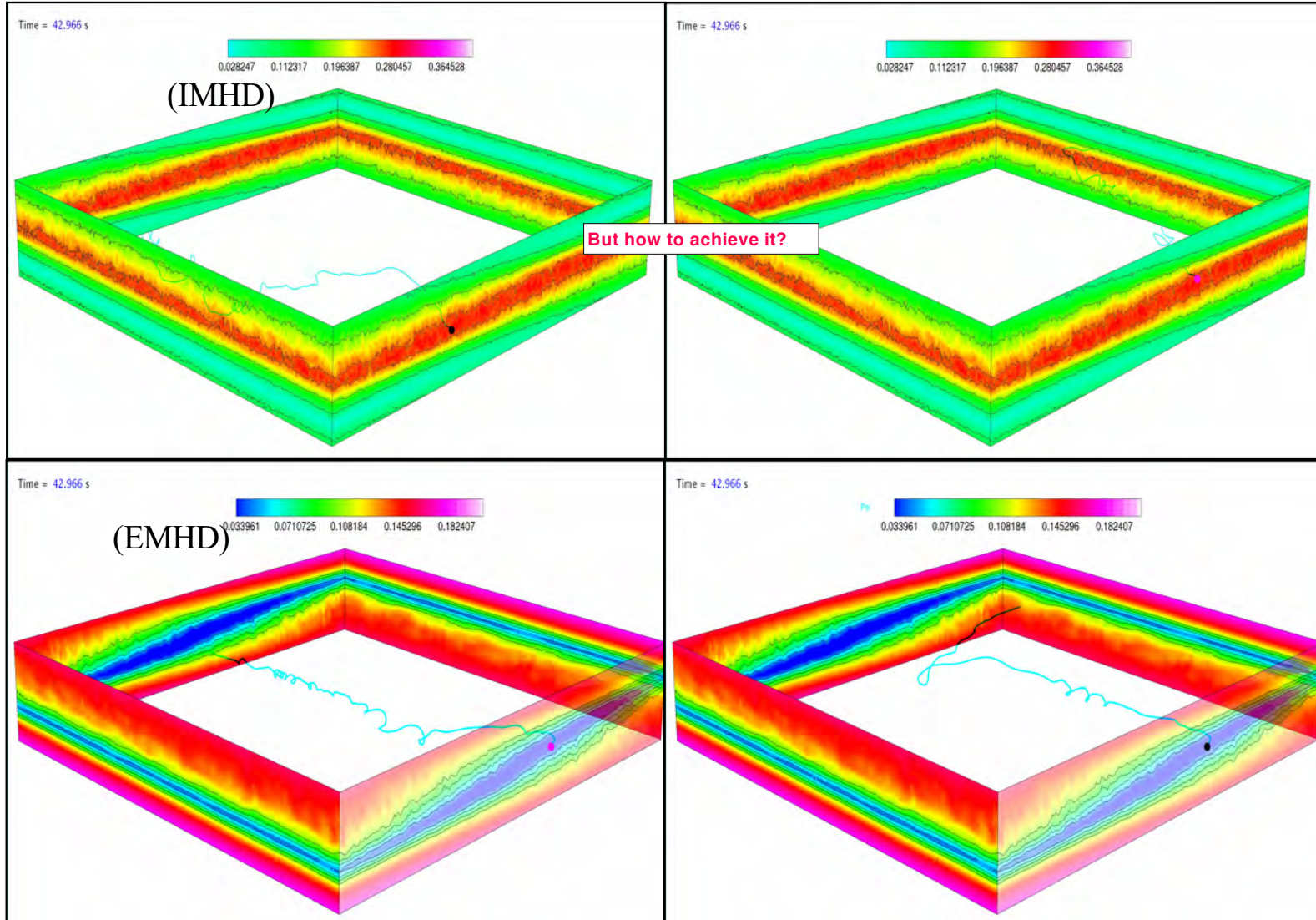
# RHPIC-LBM validation

3D turbulence acceleration model I



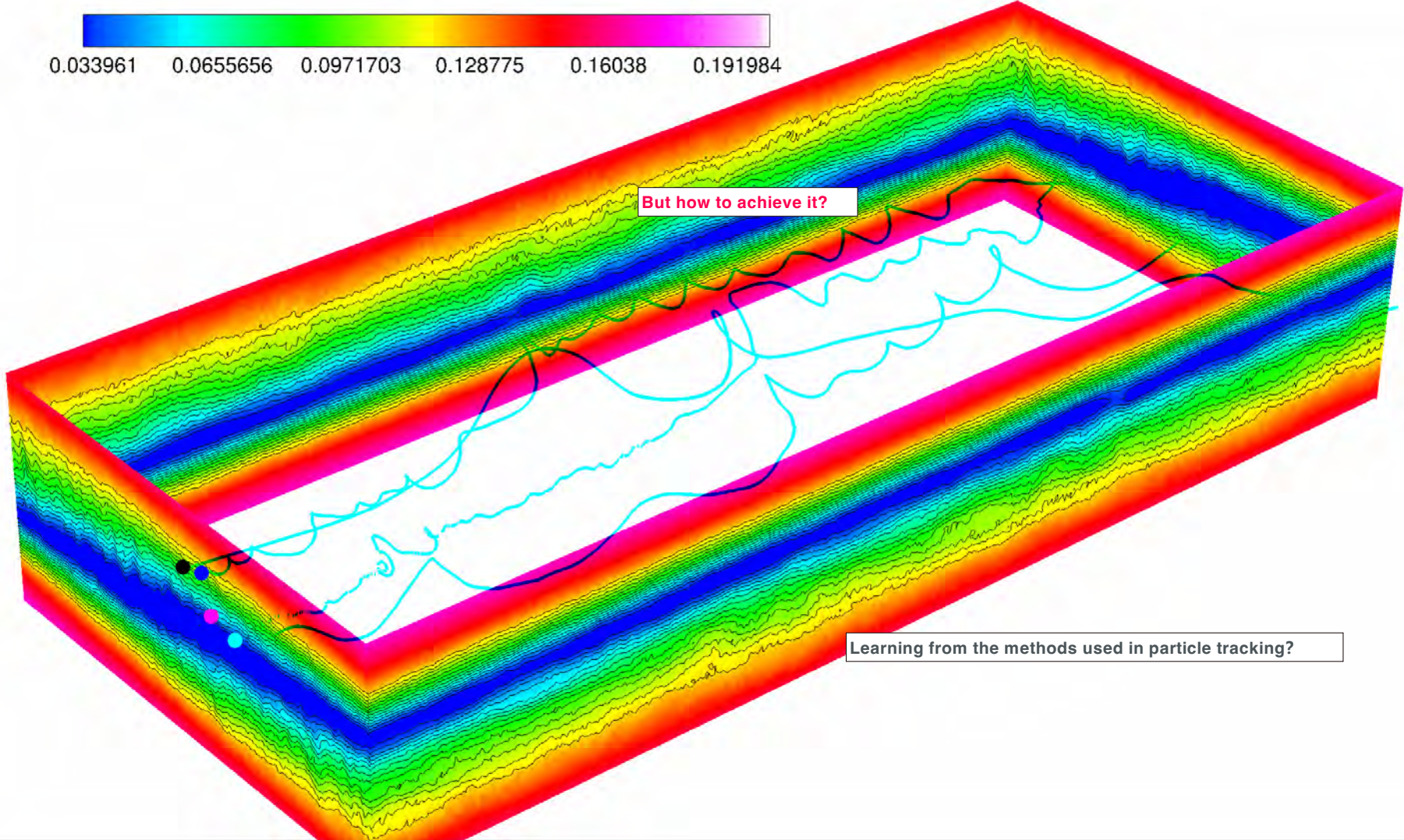
# RHPIC-LBM validation

## 3D turbulence acceleration model I



# RHPIC-LBM validation

## 3D turbulence acceleration model I

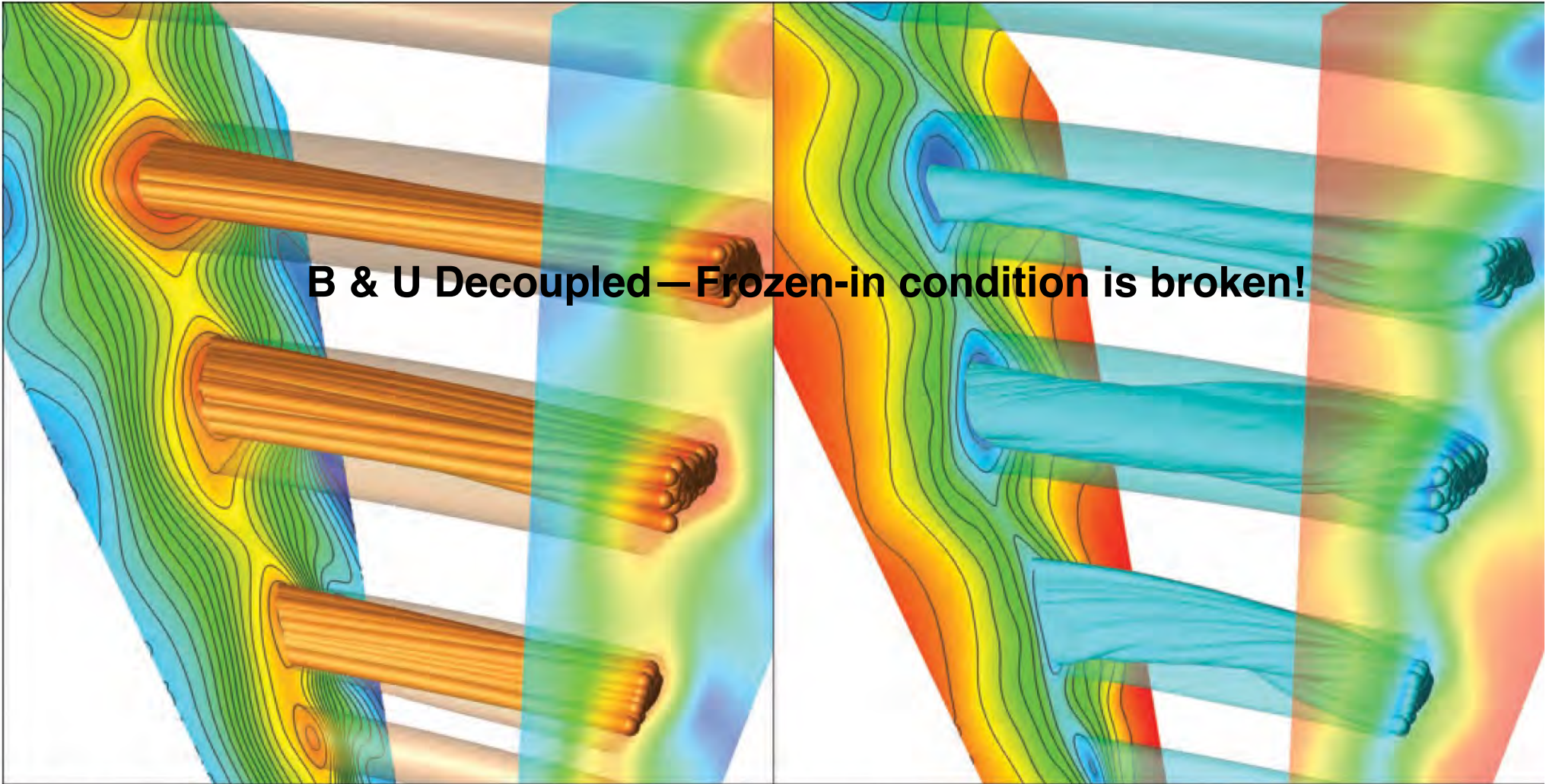


But how to achieve it?

Learning from the methods used in particle tracking?

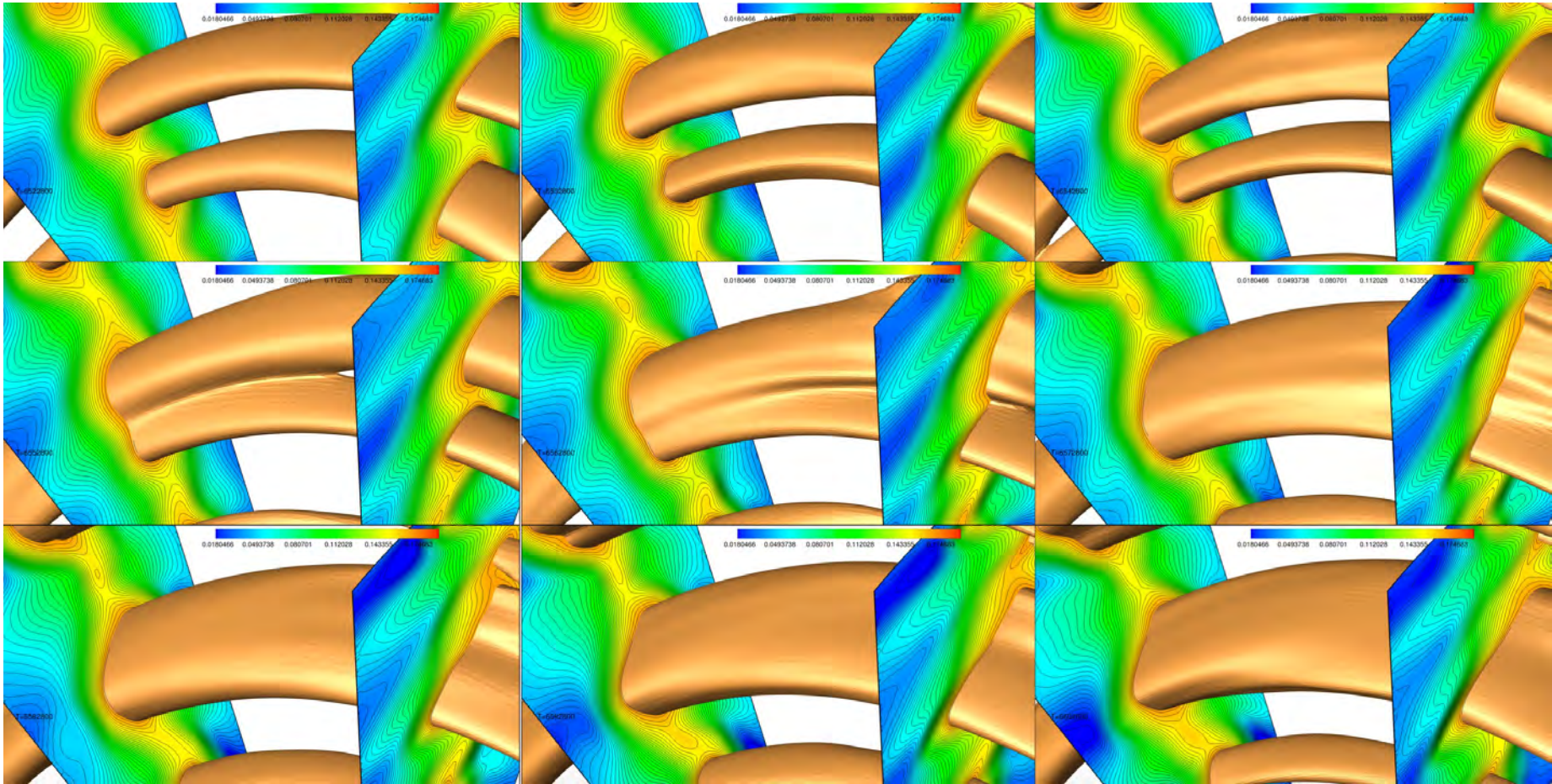
# RHPIC-LBM validation

3D turbulence acceleration model II-B&U decoupled & Instabilities



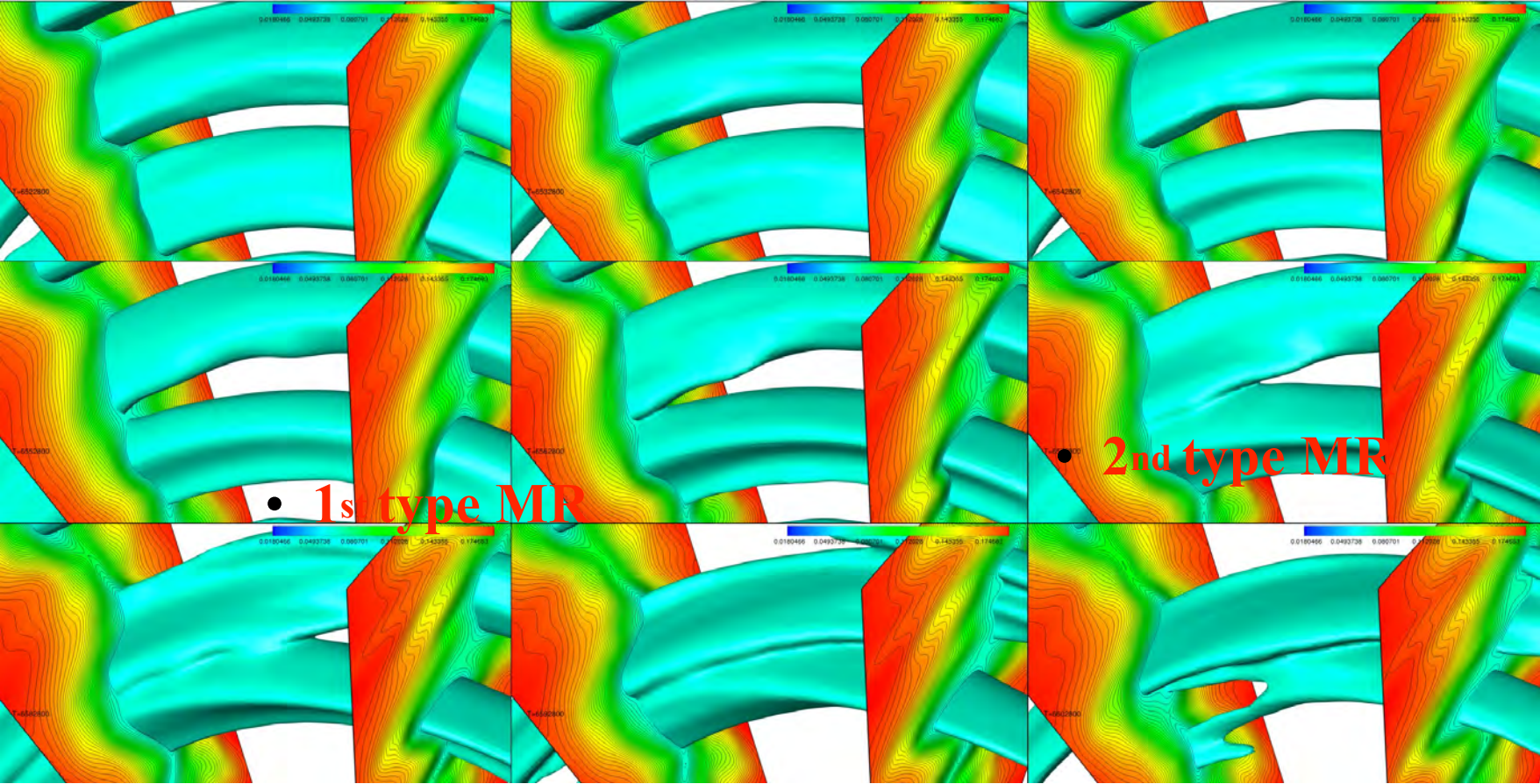
# RHPIC-LBM validation

## 3D turbulence acceleration model II-B&U decoupled & Instabilities



The cross-section of the multi-loops (MFLs/PFLs),  $d$ , is the approximately circular, ellipse or long ellipse cross-section with the equivalent diameter equal to the width between the upper and lower boundaries of the loops (MFLs/PFLs), vary from 0.6 Mm to 2 Mm (1 arcsec  $\sim$  3 arcsecs) and the corresponding peak densities from  $n$  to  $4n$

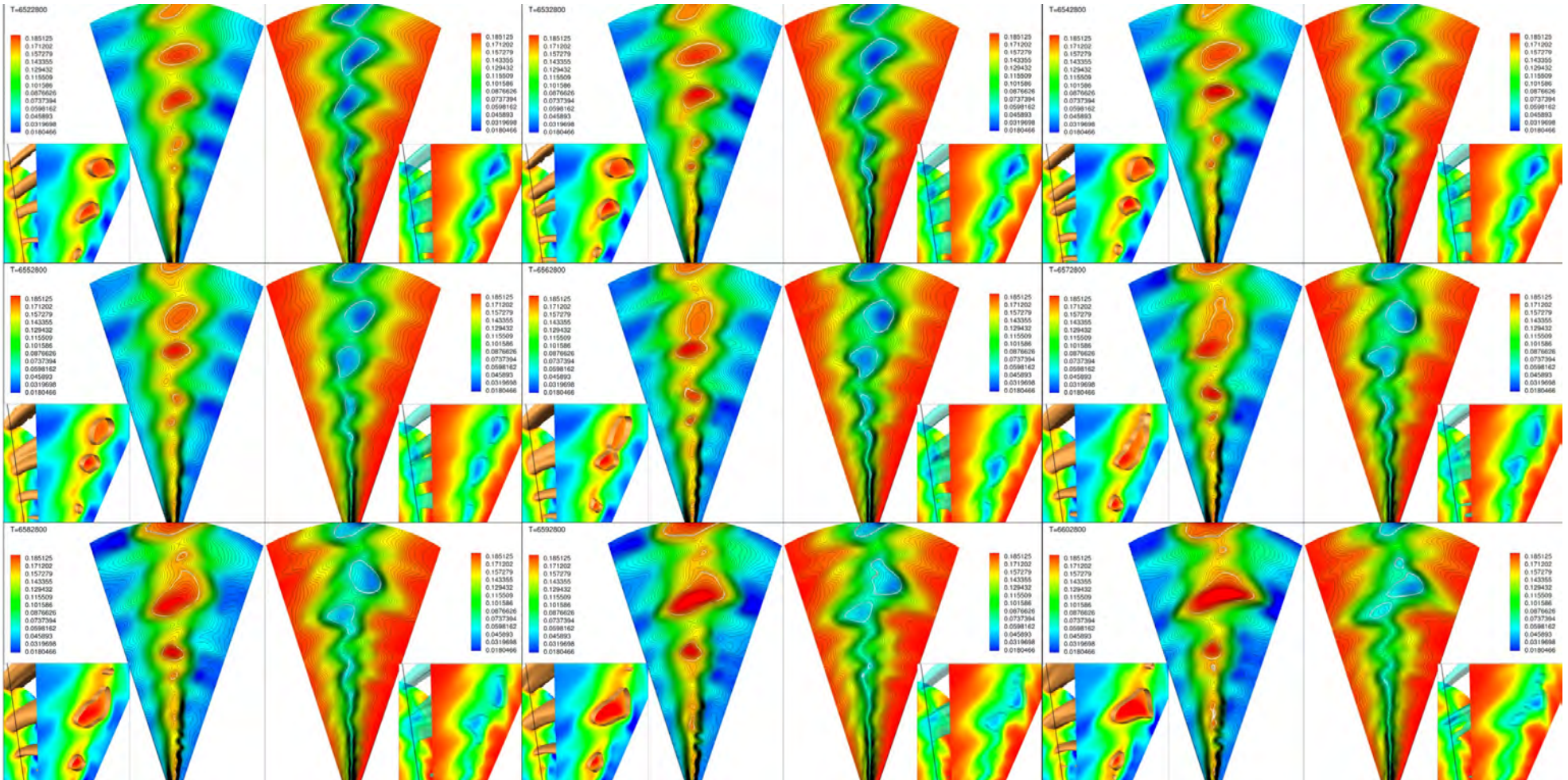
# 3D turbulence acceleration model II-B&U decoupled & Instabilities



The cross-section of the multi-loops (MFLs/PFLs),  $d$ , is the approximately circular, ellipse or long ellipse cross-section with the equivalent diameter equal to the width between the upper and lower boundaries of the loops (MFLs/PFLs), vary from 0.6 Mm to 2 Mm (1 arcsec  $\sim$  3 arcsecs) and the corresponding peak densities from  $n$  to  $4n$

# RHPIC-LBM validation 3D turbulence acceleration model II-B&U decoupled & Instabilities

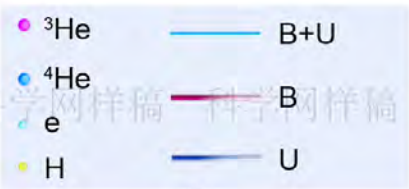
- The interaction of loops (MFLs/PFLs) decreases with height increase,



### 3. Application (Events Study)

3D turbulence acceleration model [source and origin, proton, 3He rich]

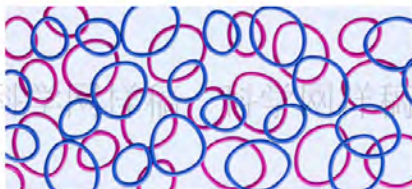
# Source and Origin



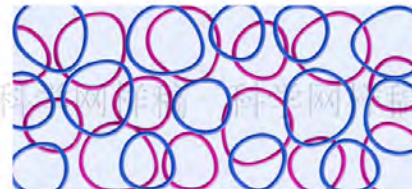
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c1



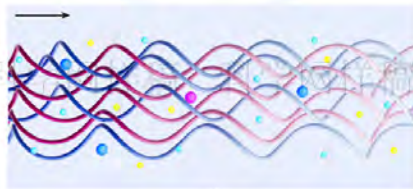
d1



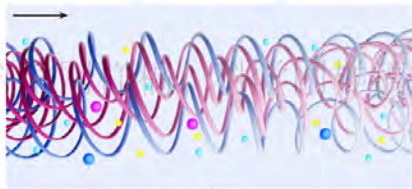
e1



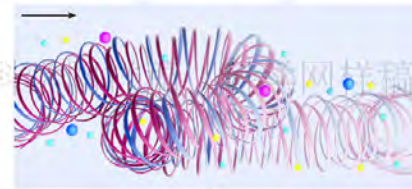
b



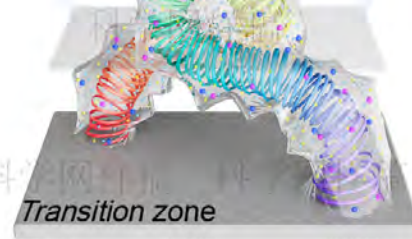
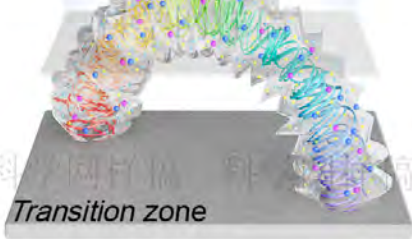
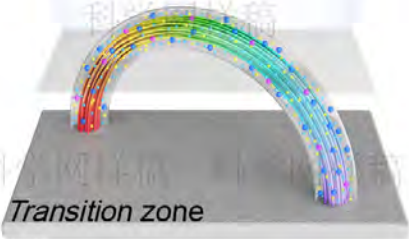
c



d

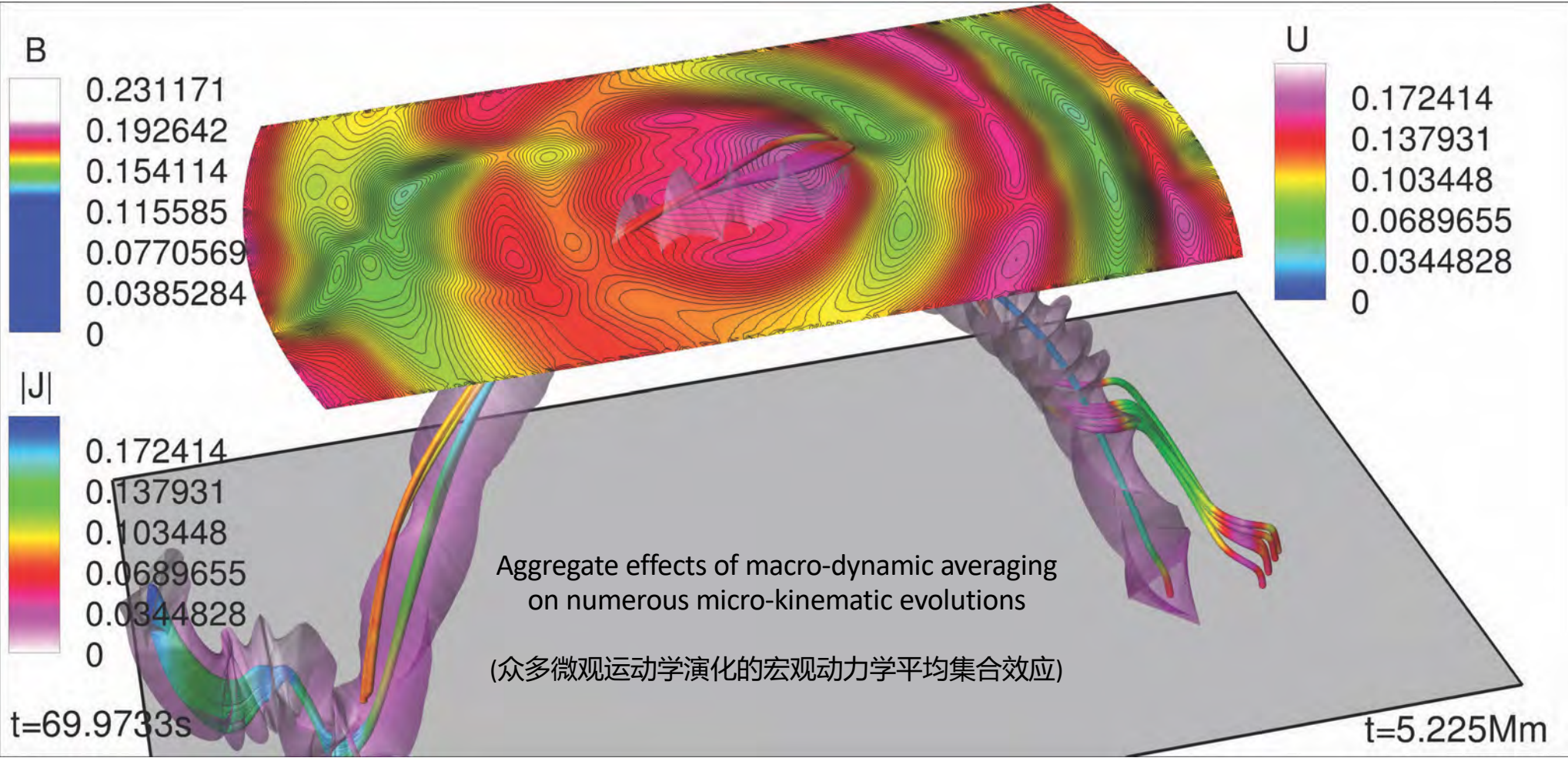


e



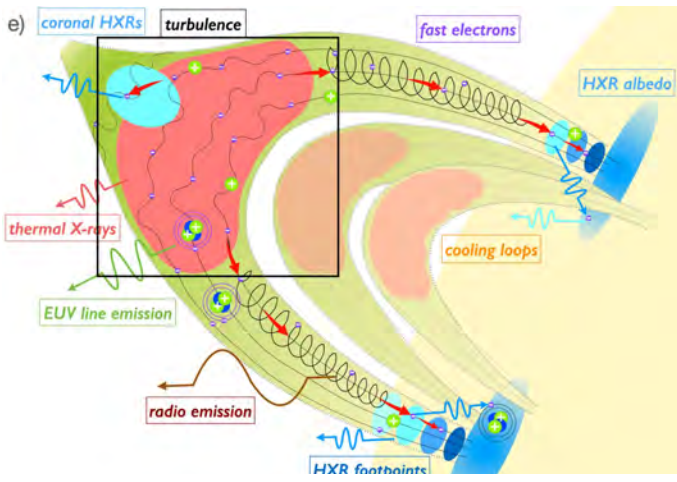
# Source and Origin

Turbulence on the top of loop (Twisted Flux Rope wrapping and unwrapping)

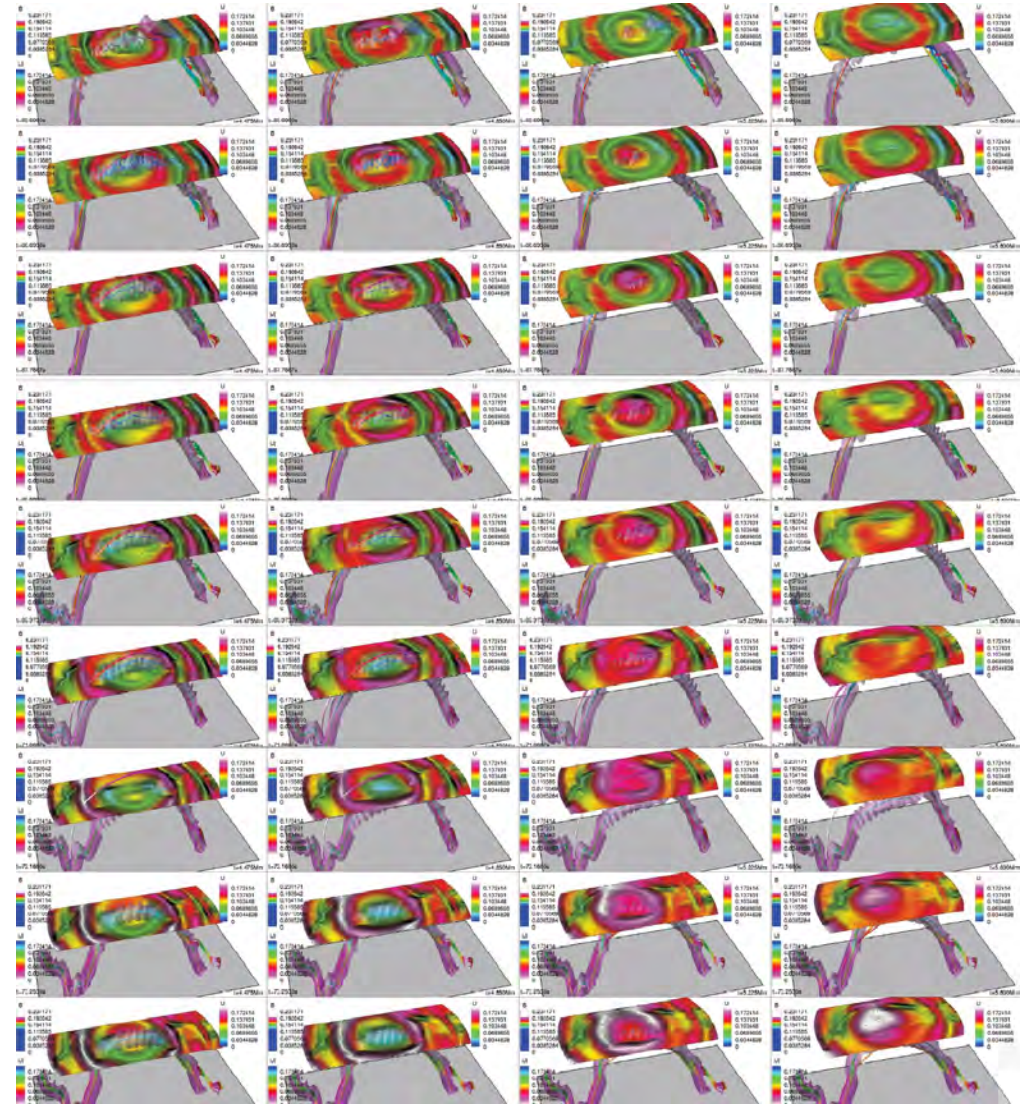
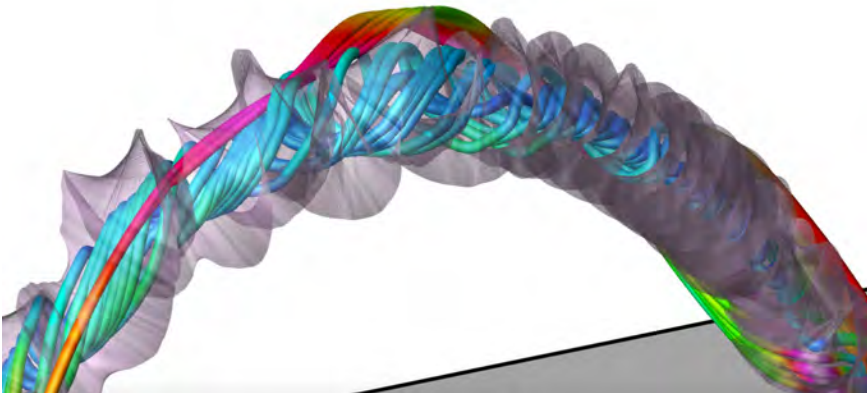


# Source and Origin

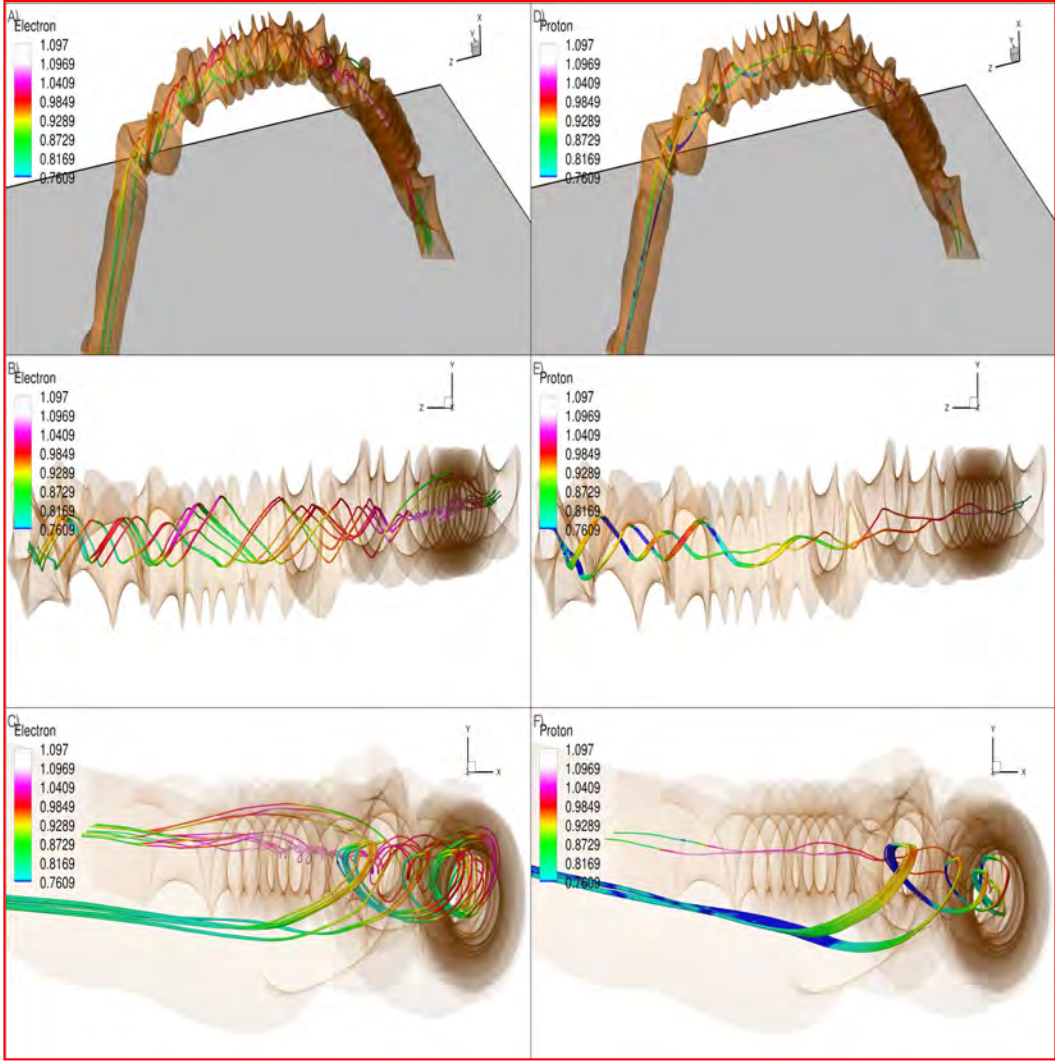
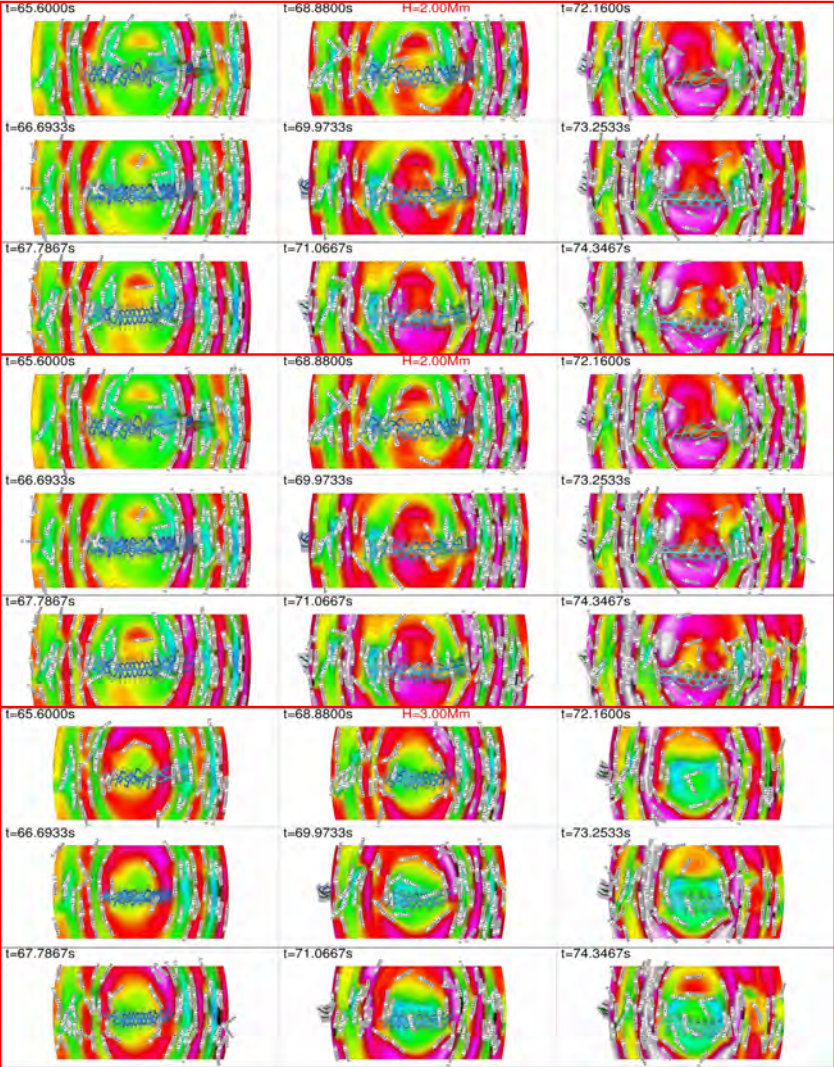
## Turbulence on the top of loop (Twisted Flux Rope wrapping and unwrapping)



Turbulence\_Twisted Flux Rope

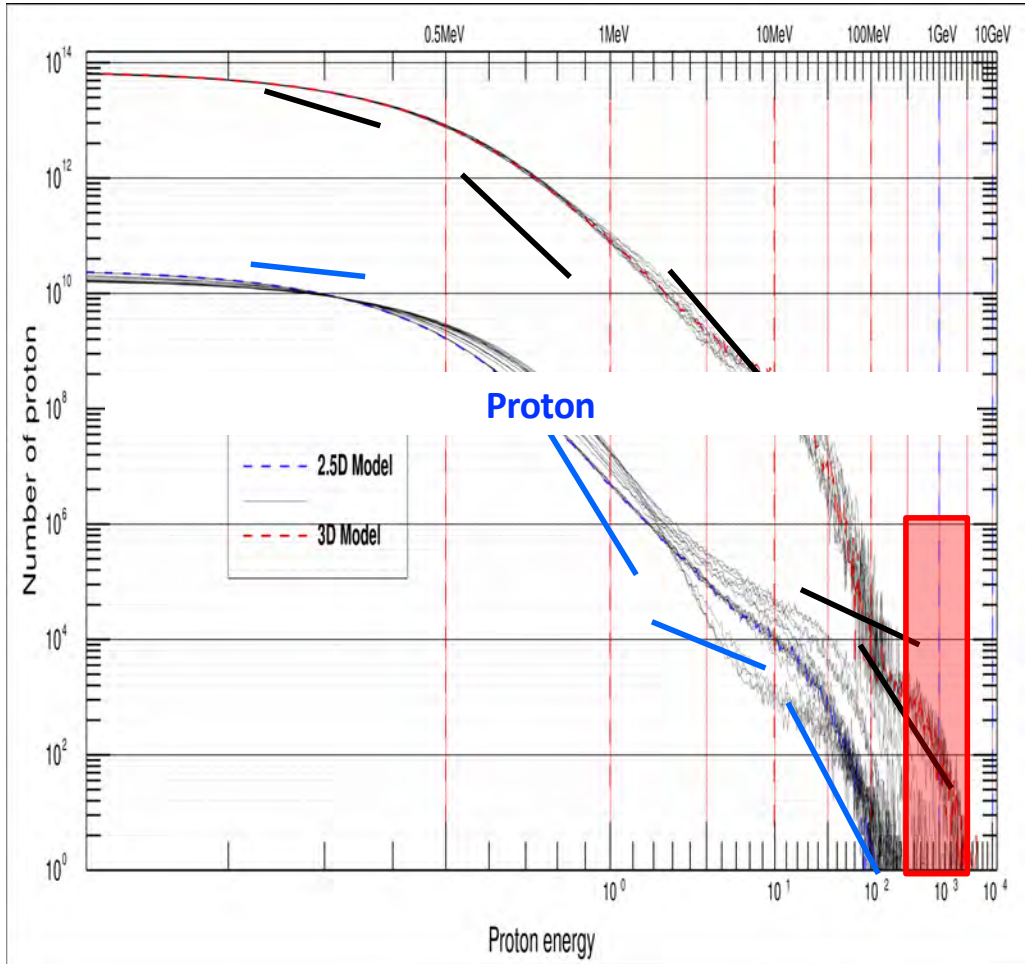
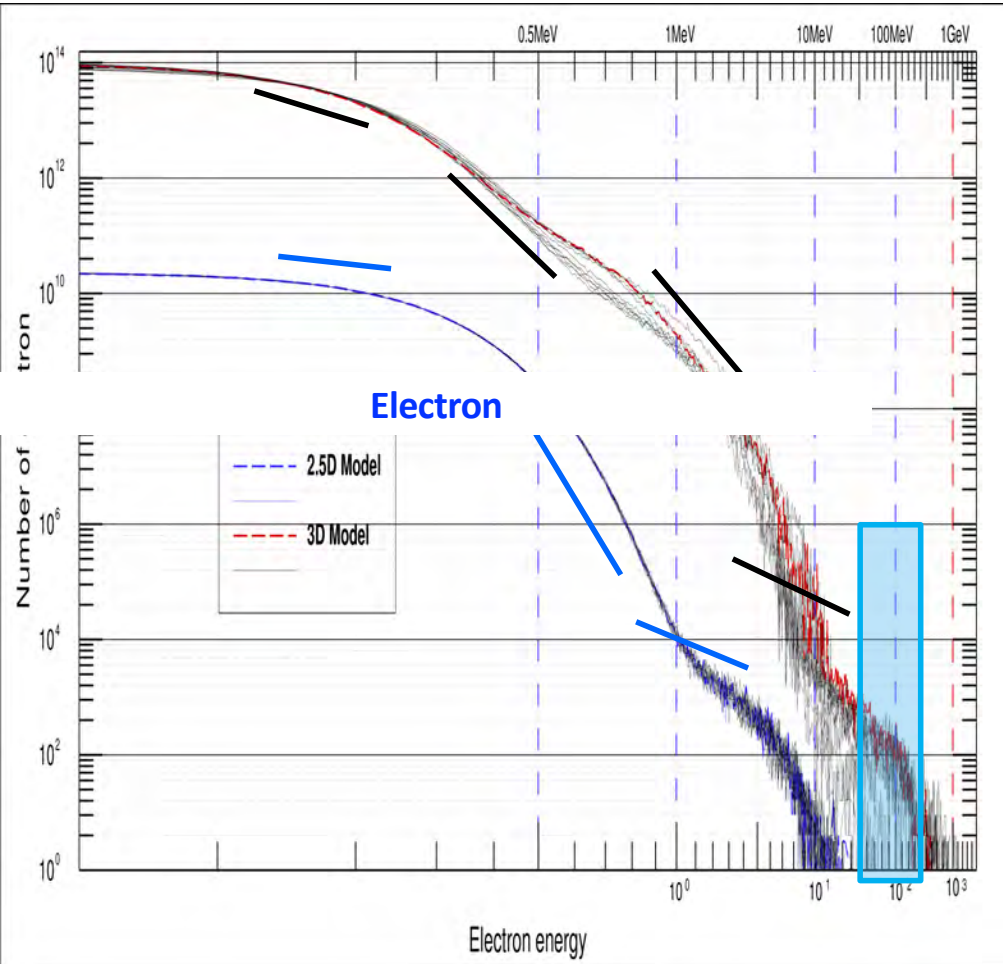


# Source and Origin 3D turbulence acceleration model-Source and origin (Proton & electron)



# Source and Origin 3D turbulence acceleration model-Source and origin (Proton & electron)

$m_p/m_e = 1836$     $m_e = 1$



**2.D vs 3D turbulence acceleration plus turbulence-induced TVD and TRD**

## **Source and Origin** 3D turbulence acceleration model-Source and origin (Proton &electron)

### **Conclusion**

The results show that protons and electrons could be efficiently accelerated simultaneously through the nonlinear resonant wave-particle interaction in the diffusion region, and the interaction of helical magnetic structure leads to efficient energization of electrons, shown as follows:

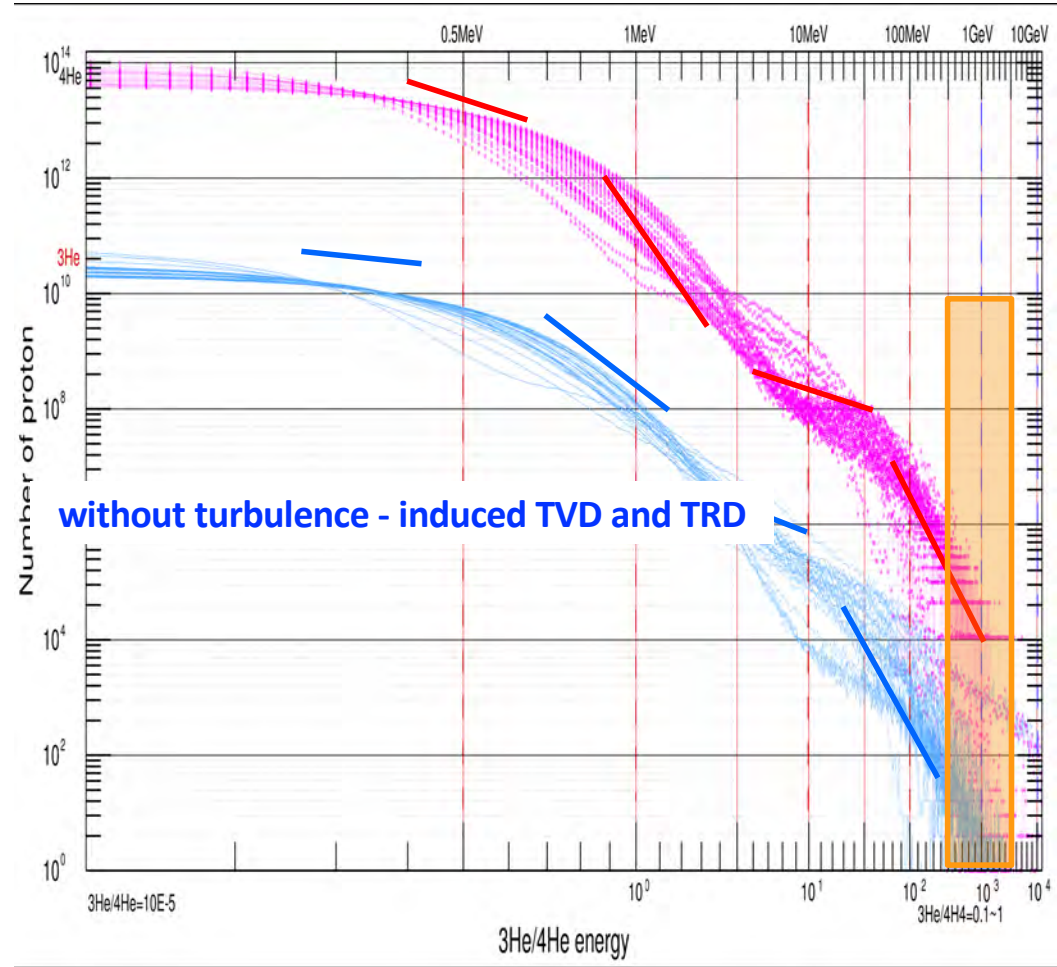
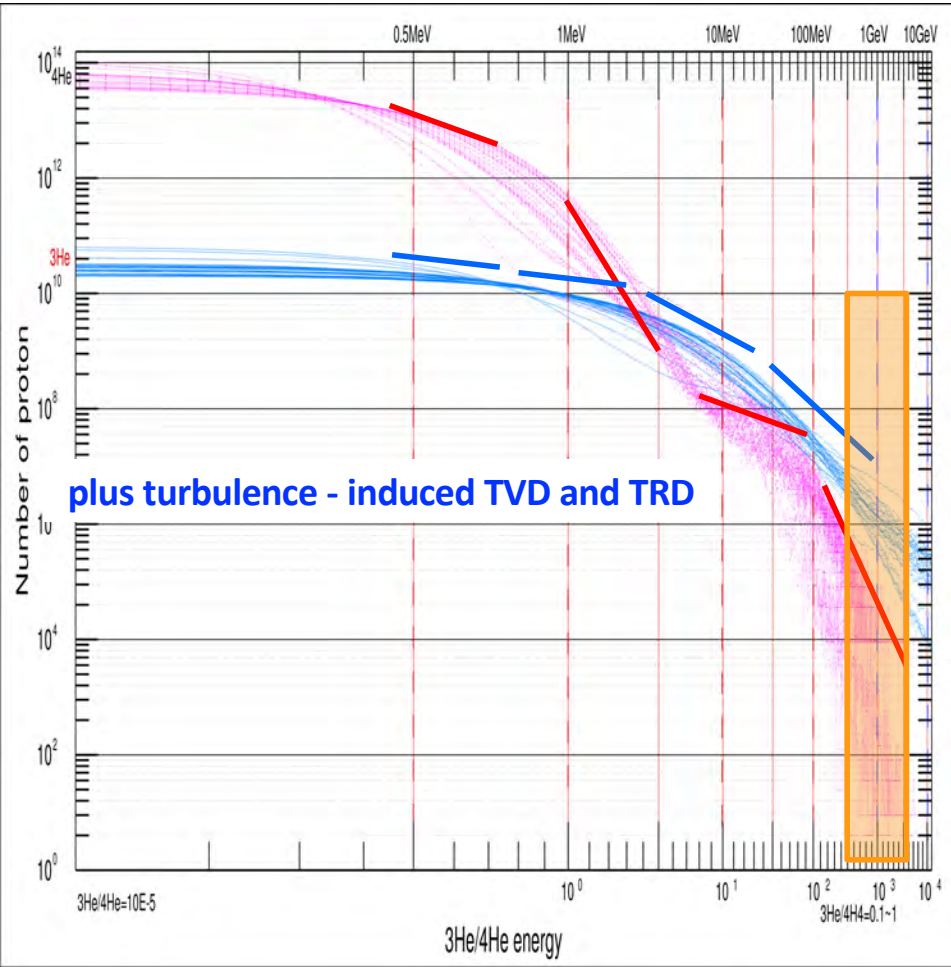
- 1. A new nonlinear resonant wave-particle acceleration mechanism in the interaction of helical magnetic structure in large-scale CME/Flare current sheets proposed in solar eruption. The original acceleration model assumed helical magnetic structure evolution in the cross-scaling coupling are separated independent process, which is essential in the helical magnetic structure coalescence and growth process.**
- 2. The multi-component acceleration mechanism and the quantitative relationship between solar flares-CME events and GeV-level SEPs events with the data-driven RHPIC-LBM algorithm and parallel code. We hope this provides a new understanding of SEP sources and origins.**
- 3. We found the Langmuir turbulence acceleration (LTA) through the nonlinear resonant wave-particle interaction by tracking the trajectories and analyzing the energy spectrum of energetic protons and electrons. LTA is an independent acceleration mechanism similar to shock acceleration. Still, it is much more efficient than shock acceleration, indicating that large-scale reconnection is a good candidate for the efficient acceleration of protons and electrons in solar eruption. In summary, we anticipate identifying the source and origin of SEPs as a critical point for understanding the relationship between SEPs acceleration and explosive energetic electrons observed in the solar flares during MHD Alfvén turbulence translates into Kinetic Alfvén turbulence progress and provides a tool for extreme space weather SEPs forecasting.**

Aggregate effects of macro-dynamic averaging on numerous micro-kinematic

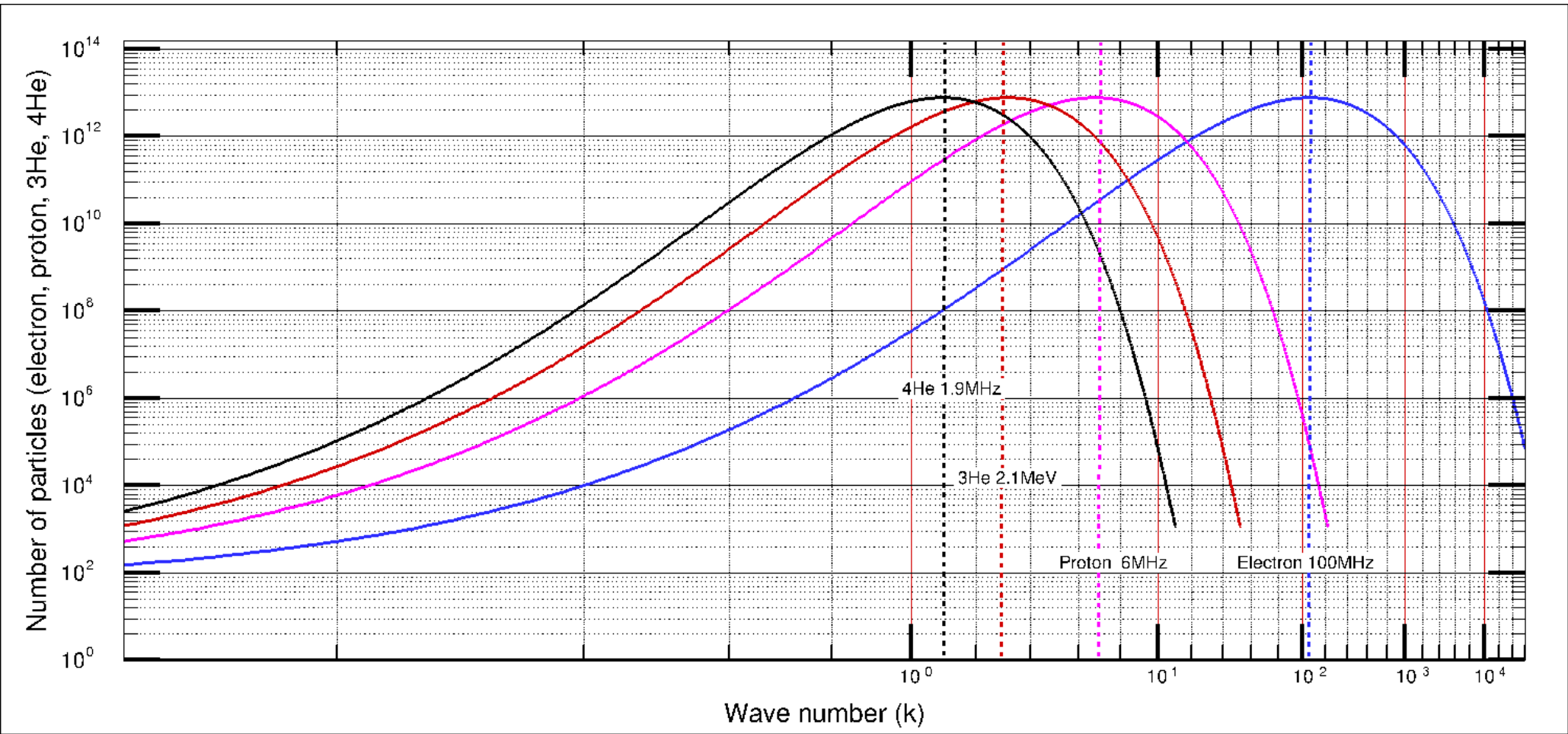
# Source and Origin

## 3D turbulence acceleration model-Source and origin (3He rich)

$$m_n/m_e = 1836 \quad m_p/m_e = 1836 \quad m_{3He} = 2m_p + m_n = 5508 \quad m_{4He} = 2m_p + 2m_n = 7344 \quad m_e = 1$$



### Wave-particle interaction turbulence acceleration



## PSP、Solar Orbiter Observational data

Electrons and Ions are accelerated out of thermal equilibrium due to magnetic reconnection event

### Electrons

1.  $10^{36}$  per sec accelerated for  $\sim 100$  s
2. Maximum energy  $\sim 100$  MeV

### Ions

1.  $10^{35}$  per sec accelerated for  $\sim 100$  s
2. Maximum energy  $\sim 10$  GeV

Nariaki V. Nitta讨论“三尺度五阶段耦合湍流”对 $^3\text{He}$ 丰度之谜如何更好解释问题

Irina Kitiashvili&Tamar Ervin,Chadi Salem,Ace Stratton, Meredith Wieber等讨论了PSP观测质子10GeV能级/电子100MeV与RHPIC-LBM模拟合作

Patrick antolin, Stephane Regnier, Yash Saneshwar, Eamon Scullion等讨论了RHPIC-LBM自主知识产权软件合作

## Source and Origin

3D turbulence acceleration model-Source and origin (3He rich)

## Conclusion

To 3He-rich ( $3\text{He}/4\text{He}$  ratio  $>0.01$ ) events showing up to 10,000-fold abundance enhancement in the pulse solar hurricanes have been a puzzle for more than 60 years since their discovery.

In this work, based on the statistical physics theoretical model combined with the filter theory of turbulence, we investigate the interaction of 3He/4He, proton, and electron particles acceleration process by the actual ratio of the proton mass to the electron mass and turbulence resistivity & viscosity item under fully coupled kinetic-dynamic continuous scale (instead of the micro-kinetic and macro-dynamic scale separated independent process) on the supercomputer with relativistic hybrid particle-in-cell & lattice-Boltzmann (RHPIC-LBM) innovative algorithm and parallel code.

1. We have discovered a new independent turbulence, the wave-particle interaction acceleration. This novel nonlinear acceleration mechanism is distinct from the current turbulence acceleration mechanism, including 1st and 2nd fermi acceleration; it is more efficient at the GeV level than the shock wave acceleration mechanism. This finding has profound implications for our understanding of space weather disasters and the safety of space orbiters.

2. Tracking the trajectories and analyzing the energy spectrum of energetic 3He/4He allows protons and electrons to be efficiently accelerated simultaneously through the nonlinear resonant wave-particle interaction in the diffusion region. Due to the different proton-to-electron mass ratios, more 3He/4He particles have reached the GeV level.

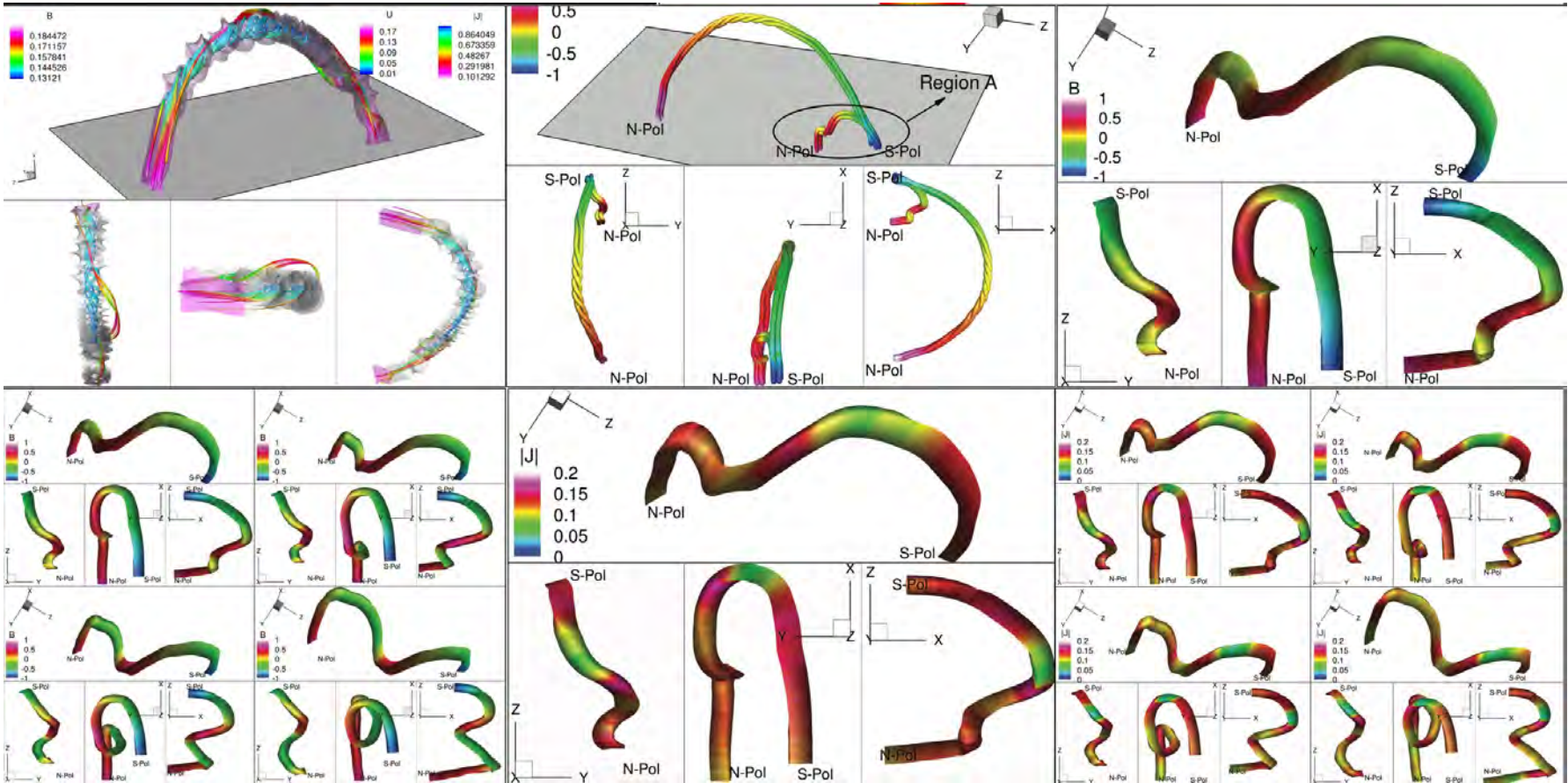
3. We found that Langmuir turbulence acceleration (LTA) through the nonlinear resonant wave-particle interaction prefers to accelerate particles with a resonance frequency of 3He, which lead is much more efficient than 4He particles.

Aggregate effects of macro-dynamic averaging on numerous micro-kinematic

## **4. Summary and Future work**

# Science point:

1. What is the difference between shock wave acceleration(turbulence acceleration;1st Fermi, and 2nd Fermi) and wave-particle interaction acceleration
2. Other heavy particles
3. Different ionization states



# The 12th International Workshop on GPU, MIC, Big Data and Cloud Computing Solutions to Multi-Physics Problems

2025CGU <http://www.cgu.org.cn/cugs/?q=node/158&page=3>, 2025.10.18-22, Chengdu



[CGU-2025 http://www.cgu.org.cn](http://www.cgu.org.cn) October 18-22 Chengdu

(2)High-Performance Computing Space-Weather/Electromagnetic-Environment Disasters

**Conveners:**Bojing Zhu,Yongbing Li,Wu Wang,Yonghui Li,Henry Tufo

The plasma in the panoramic view Sun-Earth system includes the Earth's outer core plasma that the main material general the geomagnetic filed, the upper atmosphere and the Earth's magnetosphere environment that protect the high-energy radiation and seriously affects human microwave communications, and the flare/CME plasma that the main reason for Earth's electromagnetic environment.Space-weather/electromagnetic-environment disasters can cause significant catastrophic damage to in-orbit spacecraft, astronauts, deep space exploration, and ground-based electrical facilities.it is identified as a priority area among the five key scientific themes outlined in the National Mid- and Long-term Plan for Space Science in China (2024–2050).It will be held in conjunction with the 12th International Workshop on GPU, MIC, Big Data,Cloud Computing, and AI. Conference topics:①Computational space electromagnetic environment and associated fields;②Computation space weather and associated fields;③ Other related topics.



In memory of **Prof. Dr. David Alexander Yuen**, who passed away in 2023, for his invaluable support in my academic research career.

Global Astro and Space Science Congress (GASSC2026), taking place in London, UK, from May 04–05, 2026  
Global Astro & Space Science Congress, where Astro & Space Science minds unite to shape the future. To discover groundbreaking research, connect with top professionals, and explore innovations transforming Astro & Space Science

<https://globalastrospacesciencecongress.com/core-committee>

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Bojing Zhu  
Yunnan Observatories,  
Chinese Academy of  
Sciences  
China



Doron Lancet  
Weizmann Institute of  
Science  
H Index-98  
Israel



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University of Bahrain  
Kingdom of Bahrain



Alexander Ramm  
Kansas State University  
USA



Sidy SY  
Founder, Club  
Astronomie Senegal  
Amateurs (CASA)  
South Africa



Yew Kee WONG Eric  
Hong Kong Chu Hai  
College  
Hong Kong



Eduard Babulak  
University of  
Staffordshire  
United Kingdom



Lyndon Errol Ashmore  
Independent Researcher  
UK

